

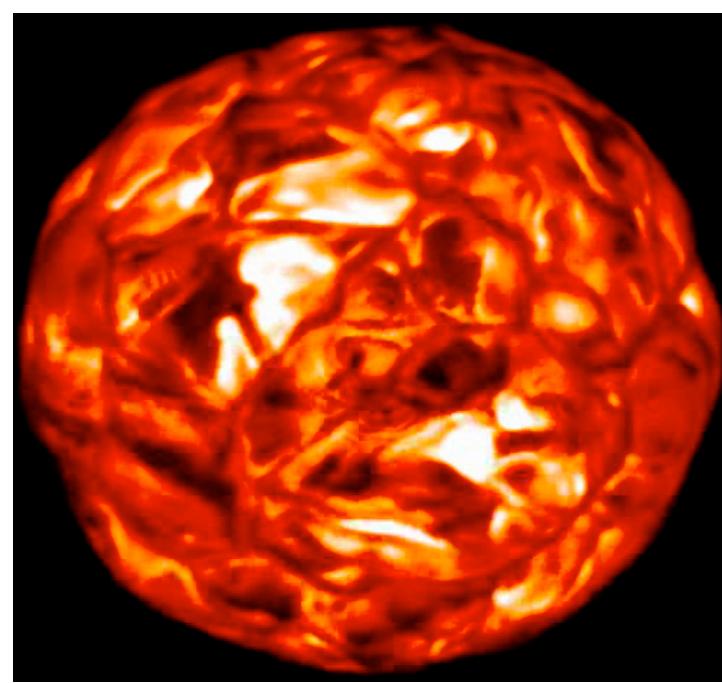
# **IMAGE-OI: an OIFITS extension for image reconstruction algorithms**

## **Application to Olmaging**

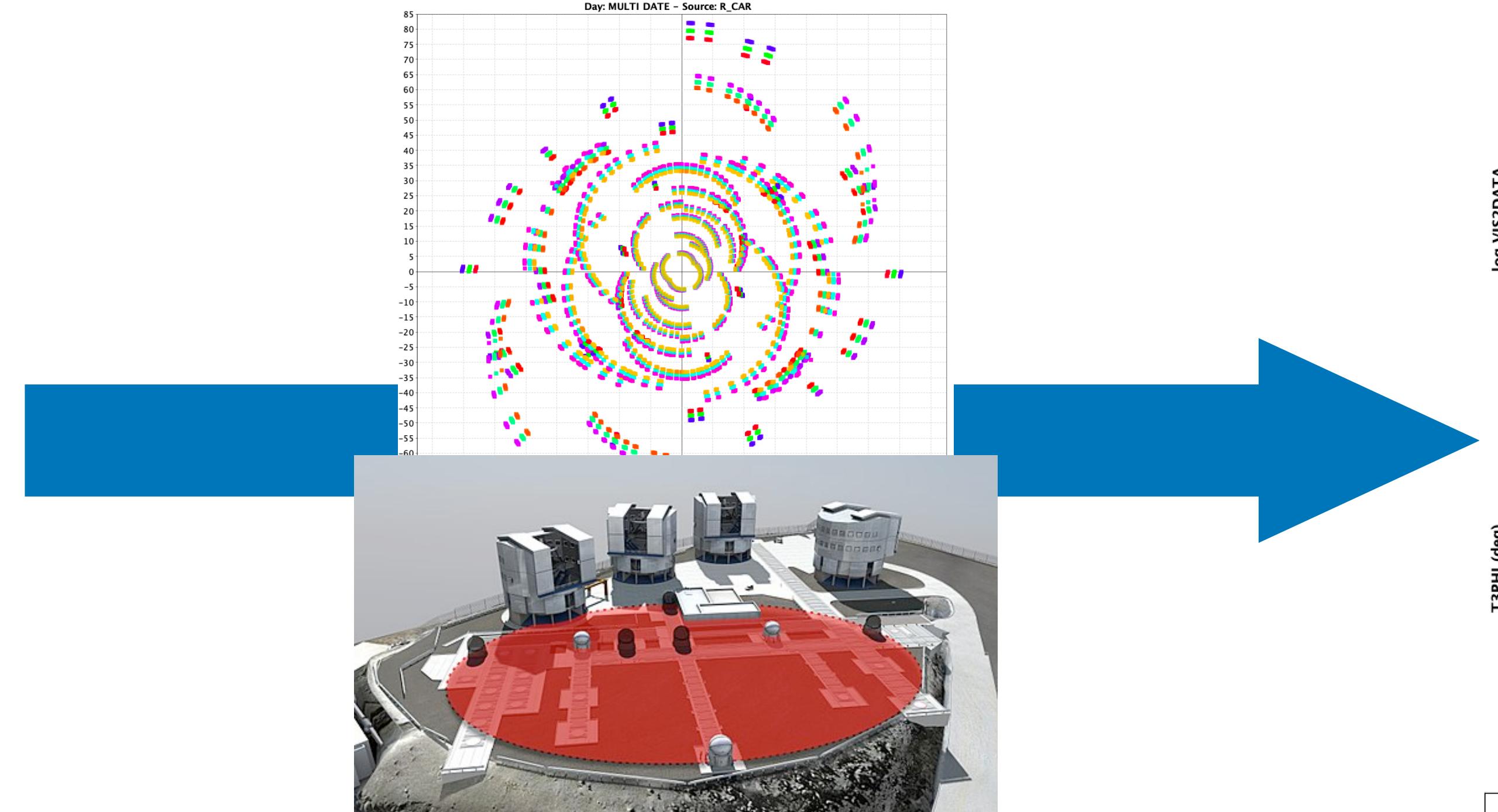
**JMMC**

**F. Soulez, L. Bourgès, A. Kaszczyc, G. Mella, M. Pratoussy, G. Duvert, J. Kluska, E. Thiébaut and J. Young**

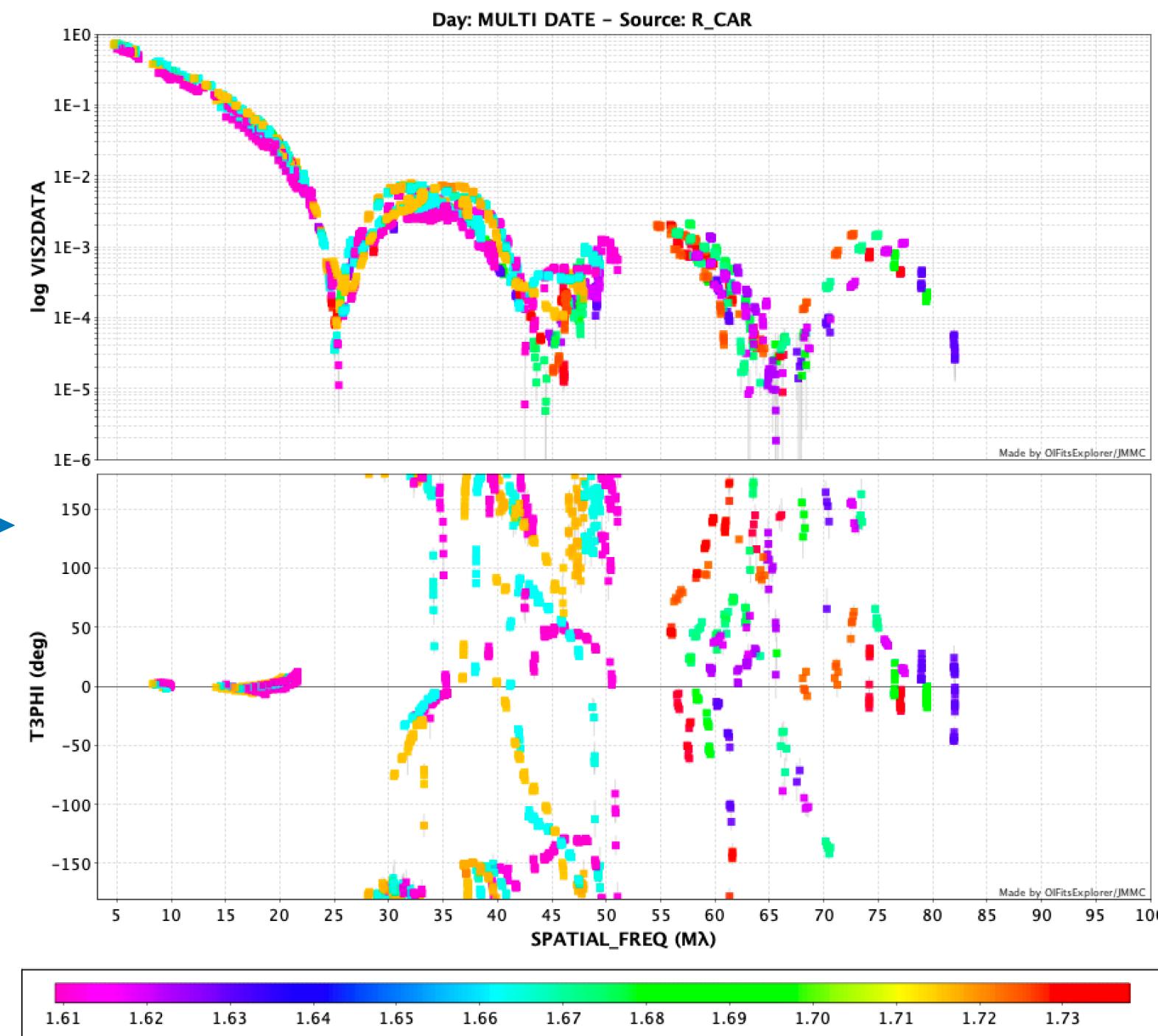
# Image reconstruction in interferometry



Object



Interferometric instrument



Measurements



Image reconstruction

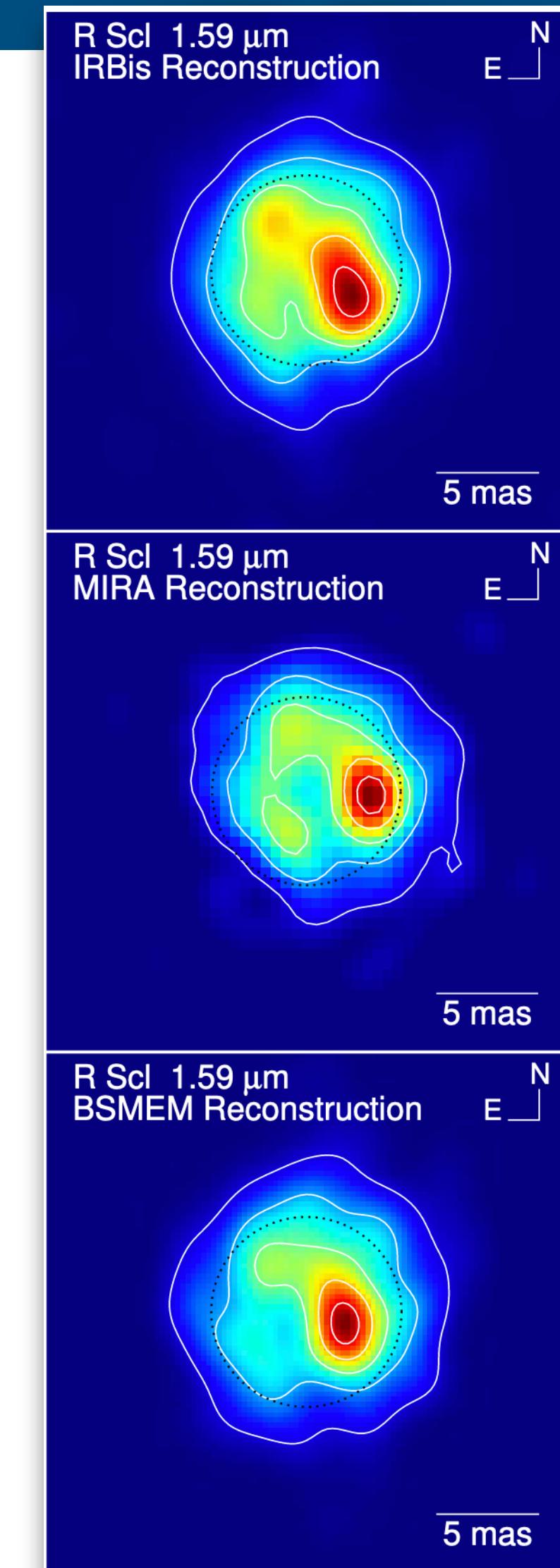
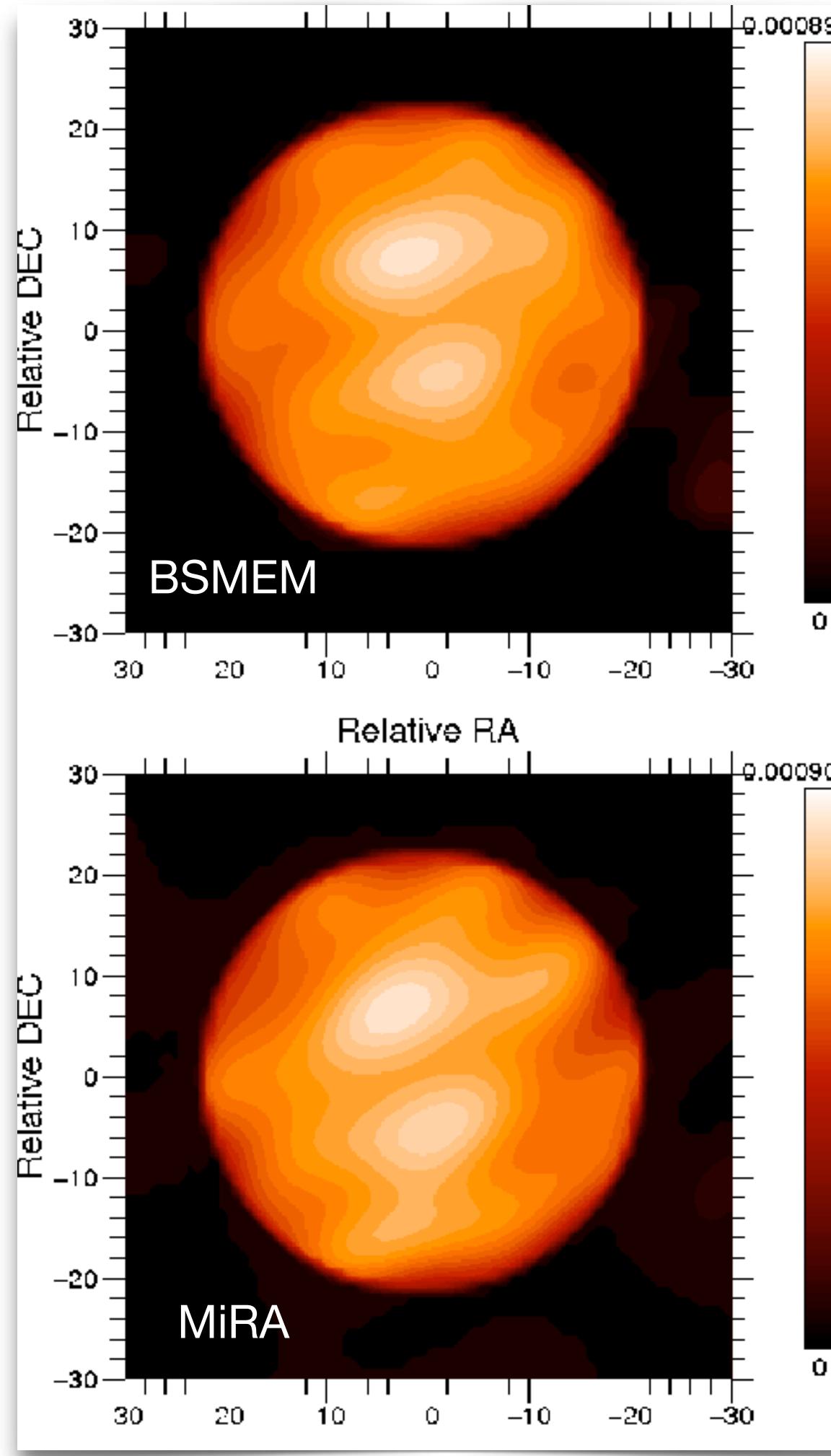
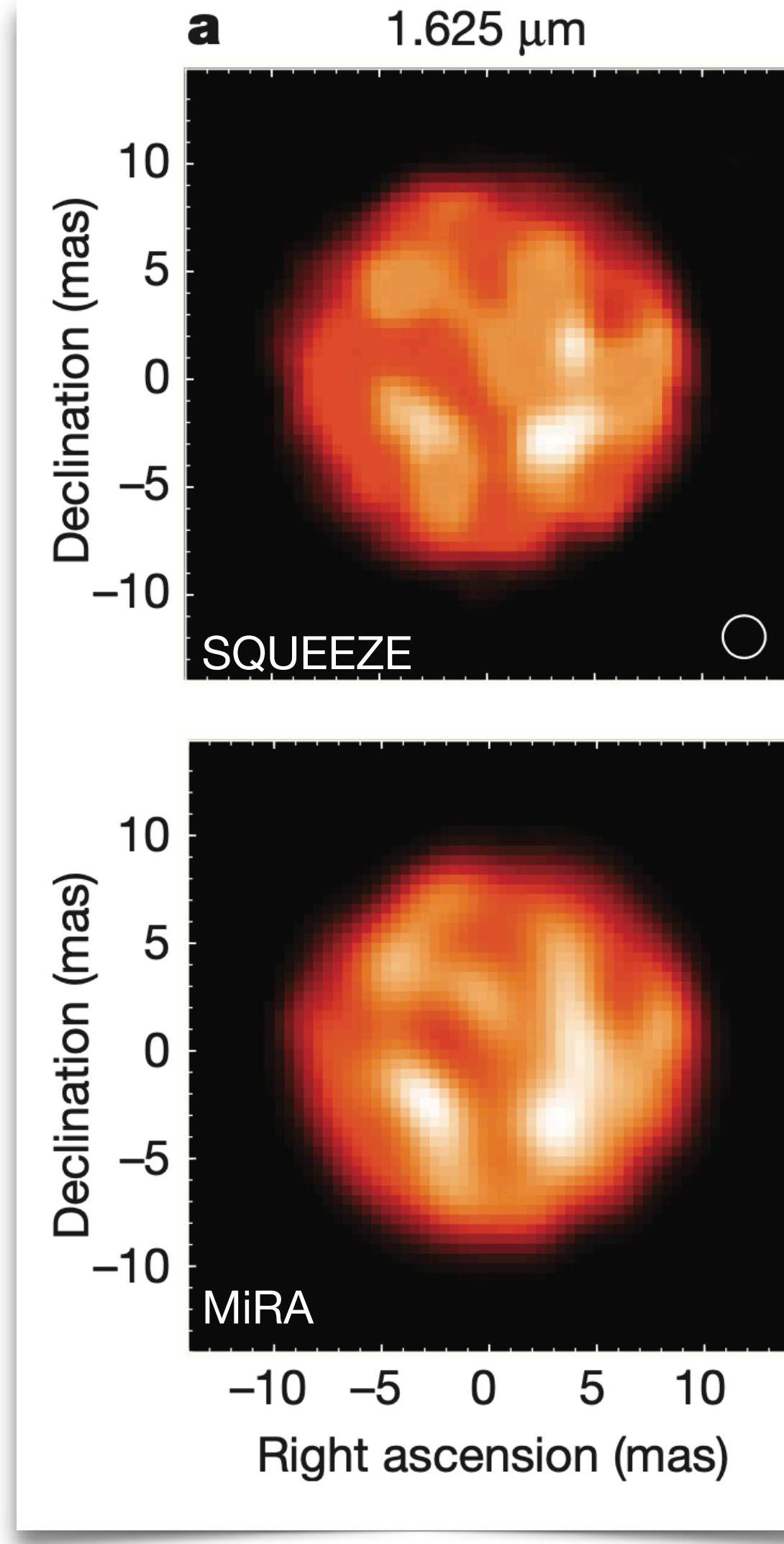
# Image reconstruction softwares

- **BSMEM** (Buscher et al, 1994)
- **MACIM** (Ireland et al, 2006)
- **MiRA** (Thiébaut, 2008)
- **WISARD** (Mugnier et al 2008)
- **SQUEEZE** (Baron et al, 2010)
- **IRBIS** (Hoffman et al, 2014)
- **SPARCO** (Kluska et al, 2014)
- **ORGANIC** (Claes et al 2020)
- **GR<sup>R</sup>** (GRAVITY col., 2022)

# Image reconstruction softwares

● <b>BSMEM</b> (Buscher et al, 1994)	C
● <b>MACIM</b> (Ireland et al, 2006)	C
● <b>MiRA</b> (Thiébaut, 2008)	yorick
● <b>WISARD</b> (Mugnier et al 2008)	IDL
● <b>SQUEEZE</b> (Baron et al, 2010)	C
● <b>IRBIS</b> (Hoffman et al, 2014)	C
● <b>SPARCO</b> (Kluska et al, 2014)	yorick or C
● <b>ORGANIC</b> (Claes et al 2020)	python
● <b>GR</b> (GRAVITY col., 2022)	python

# Needs: comparing reconstructions softwares



$\pi^1$  Gruis (Paladini, 2018)

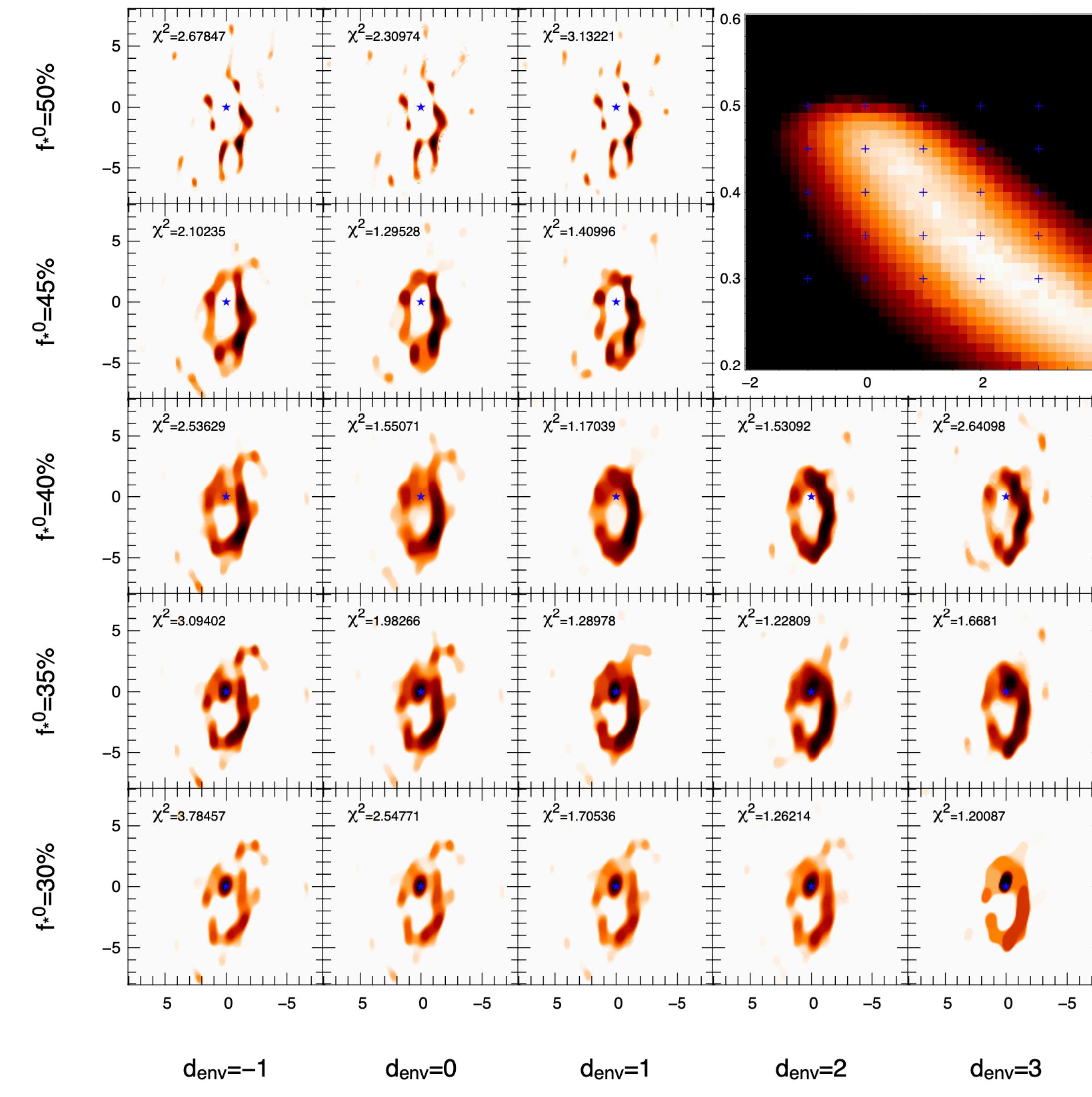
Betelgeuse (Haubois, 2009)

R Sculptoris (Wittkowski, 2017)

# Needs: benchmarking parameters

## Comparing results with different

- Priors
- Parameters
- Initialization



# Reconstruction algorithm in a nutshell

## Input:

- Initial image
  - size
  - sampling
- Data
  - selection parameters
- Parameters
  - priors
  - hyper-parameters
  - number of iterations
  - ...



## Output:

- Reconstructed image
- Model of the data
- Output parameters
  - Chi-square
  - cost function value
  - ...

# Image-OI: an OIFITS extension for image reconstruction algorithms

## Reconstructed image

Index	Extension	Type
0	Primary	Image
1	OI_TARGET	Binary
2	OI_ARRAY	Binary
3	OI_WAVELENGTH	Binary
4	OI_VIS2	Binary
5	OI_T3	Binary
6	IMAGE-OI INITIAL IMAGE	Image
7	IMAGE-OI INPUT PARAM	Binary
8	IMAGE-OI OUTPUT PARAM	Binary

## OIFITS data

## Initial image

## Input parameters

## Output parameters

# Input parameters

Input parameters stored in the header of ‘IMAGE-OI INPUT PARAM’ HDU.

Data Selection (mandatory)		
Keyword	Type	Description
TARGET	string	Identifier of the target object to reconstruct
WAVE_MIN	real	Minimum wavelength to select (in meters)
WAVE_MAX	real	Maximum wavelength to select (in meters)
USE_VIS	string	Complex visibility data to consider if any <sup>†</sup>
USE_VIS2	logical	Use squared visibility data if any
USE_T3	string	Bispectrum data to consider if any <sup>†</sup>

<sup>†</sup> value can be: ‘NONE’, ‘ALL’, ‘AMP’ or ‘PHI’.

Algorithm Settings		
Keyword	Type	Description
INIT_IMG	string	Identifier of the initial image
MAXITER	integer	Maximum number of iterations to run
RGL_NAME	string	Name of the regularization method
AUTO_WGT	logical	Automatic regularization weight
RGL_WGT	real	Weight of the regularization
RGL_PRIO	string	Identifier of the HDU with the prior image
FLUX	real	Assumed total flux (1 is the default)
FLUXERR	real	Error bar for the total flux (0 means strict constraint)
HDUPREFIX	string	Prefix that specifies the leading text to use in the HDUNAME of the final image

More complex inputs can be stored in the HDU itself

# Initial image

Initial image set the field of view and the pixel size of the reconstruction

Image Parameters		
Keyword	Type	Description
HDUNAME	string	Unique name for the image within the FITS file
<del>NAXIS1</del>	integer	First dimension of the image
<del>NAXIS2</del>	integer	Second dimension of the image
CTYPE1	string	Coordinate name 'RA---TAN' for 1st axis
CTYPE2	string	Coordinate name 'DEC--TAN' for 2nd axis
<del>CDELT<i>i</i></del>	real	Physical increment along <i>i</i> -th dimension of the image (for <i>i</i> = 1 or 2)
CUNIT <i>i</i>	string	Physical units for CDELT <i>i</i> and CRVAL <i>i</i> ; defaults to 'deg' if omitted
CRPIX <i>i</i>	real	Index of reference pixel along <i>i</i> -th dimension (for <i>i</i> = 1 or 2); defaults to the geometric center of the field of view if omitted
CRVAL <i>i</i>	real	Physical coordinate of reference pixel along <i>i</i> -th di- mension (for <i>i</i> = 1 or 2) and relative to the center of the field of view; defaults to 0 if omitted

Can be a simple shape or a  
more complex object  
generated by model fitting tool

# Reconstructed image and output parameters

Reconstructed image is stored in the primary HDU

Most FITS viewer opens the primary by default

# Reconstructed image and output parameters

Reconstructed image is stored in the primary HDU

Most FITS viewer opens the primary by default

Output parameters are stored in the header of ‘IMAGE-OI OUTPUT PARAM’ HDU

Algorithm Results		
Keyword	Type	Description
LAST_IMG	string	Identifier of the final image
NITER	integer	Total iterations done in the current program run
CHISQ	real	Reduced chi-squared
FPRIOR	real	Regularization penalty
FLUX	real	Total image flux
PROCSOFT	string	Software name and version number
CONVERGE	boolean	Set to ’T’ if the algorithm stopped because it has converged

# Modeled data

Extra column next to the data containing the model and optionally its errors

New columns in OI_VIS tables		
Label	Format	Description
NS_MODEL_VISAMP	D(NWAVE)	Model of the visibility amplitude
NS_MODEL_VISAMPERR	D(NWAVE)	Model of the error in visibility amplitude (optional)
NS_MODEL_VISPHI	D(NWAVE)	Model of the visibility phase in degrees
NS_MODEL_VISPHIERR	D(NWAVE)	Model of the error in visibility phase in degrees (optional)

New columns in OI_VIS2 tables		
Label	Format	Description
NS_MODEL_VIS2	D(NWAVE)	Model of the squared visibility
NS_MODEL_VIS2ERR	D(NWAVE)	Model of the error in squared visibility (optional)

New columns in OI_T3 tables		
Label	Format	Description
NS_MODEL_T3AMP	D(NWAVE)	Model of the triple-product amplitude
NS_MODEL_T3AMPERR	D(NWAVE)	Model of the error in triple-product amplitude (optional)
NS_MODEL_T3PHI	D(NWAVE)	Model of the triple-product phase in degrees
NS_MODEL_T3PHIERR	D(NWAVE)	Model of the error in triple-product phase in degrees (optional)

# Image-OI: a common extension to connect them all

## A unified way to call reconstruction softwares

- Every information in a single OI-FITS file
- Simpler interaction with softwares
- Reproducible results

The screenshot shows a GitHub repository page for `JMMC-OpenDev/OI-Imaging-JRA`. The repository has 6 issues, 6 pull requests, and 44 commits. The `Code` tab is selected. The `README.md` file contains the following content:

```
OI-Interface

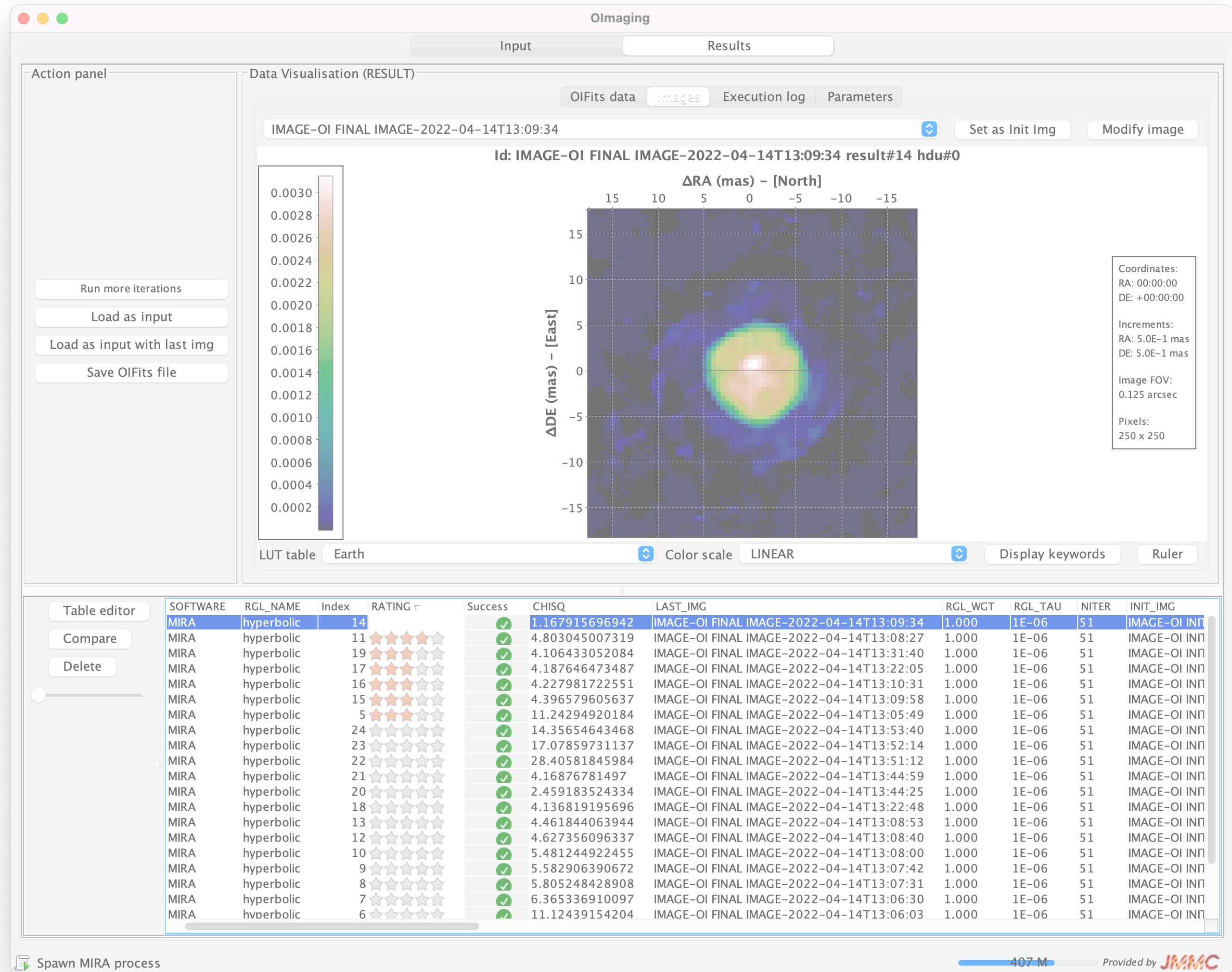
Design and specification of an interface to image reconstruction and model fitting from optical interferometric data.

Contents

• Unified Image Reconstruction Description exploits the inverse problem framework to present the general principles of image reconstruction from interferometric data.
• Interface to Image Reconstruction is a draft document giving the specifications for a graphical user interface to control image reconstruction algorithms.
• Directory doc contains the sources of the various documents.
```

On the right side of the page, there are sections for **About**, **Releases**, **Packages**, and **Contributors**.

## A single graphical user interface for image reconstruction algorithms



The diagram illustrates a dual-telescope optical interferometer setup. Two telescopes, Telescope #1 and Telescope #2, collect light from an object and a phase reference star. The light paths are shown in red and blue, converging at a star separator. A deformable mirror is used to correct the wavefronts. A laser provides a wavefront reference. The light then passes through a tip-tilt pupil and a beam combiner instrument. The combined light is directed through a delay line and a metrology sensor. The resulting image is shown in a 2° FoV. The diagram also includes labels for 'IP wavefront' and 'dOPD'. A green star indicates the 'Wavefront reference' star. A DOI link (DOI 10.1051/0004-6361/201730838, copyright ESO 2017) is provided.

**A new tool for GRAVITY-Wide**

Phase reference  
Object  
Telescope #1  
Starlight  
Metrology  
Deformable mirror  
IP wavefront  
Star separator  
Laser  
Tip-tilt Pupil  
Read more >  
2° FoV  
Metrology sensor  
Delay line  
Beam combiner instrument  
dOPD  
Read more >

DOI 10.1051/0004-6361/201730838,  
copyright ESO 2017

### The Mission

The JMMC is the french center for optical interferometry. It aims at providing support for the users of the stellar interferometers currently in operation. This support is possible thanks to the development of efficient and user friendly tools for preparing the observations, analysing the data or archiving the results. The tools are accessible through the web site and linked to a "Face to Face" help, especially for the preparation of observations, the PIONIER, GRAVITY and MATISSE data reduction, and the data analysis.

[Who are we? >](#)

### JMMC User Support

The JMMC is committed to provide support to the users of the VLTI and other interferometers. For this purpose, a single [contact e-mail address](#) has been created. You can also fill the dedicated feedback form. Access by click on "Read more".

[Read more >](#)

# JMMC Service overview



V L T I



C H A R A

SUV (VLTI Center):

+ User Support

+ Training

JSDC2  
JMDC

CDS Catalogs

Two screenshots of astronomical catalog databases. The top one is VizieR showing a search interface for stellar diameters. The bottom one is JSDC2/JMDC showing a search interface for JMMC Stellar Diameters Catalogue - JSDC Version 2.

SearchCal

A screenshot of the SearchCal software interface. It shows query parameters for a science object (Eta Tau) and an instrumental configuration (VLT, AMBER). The main window displays a plot of intensity vs wavelength.

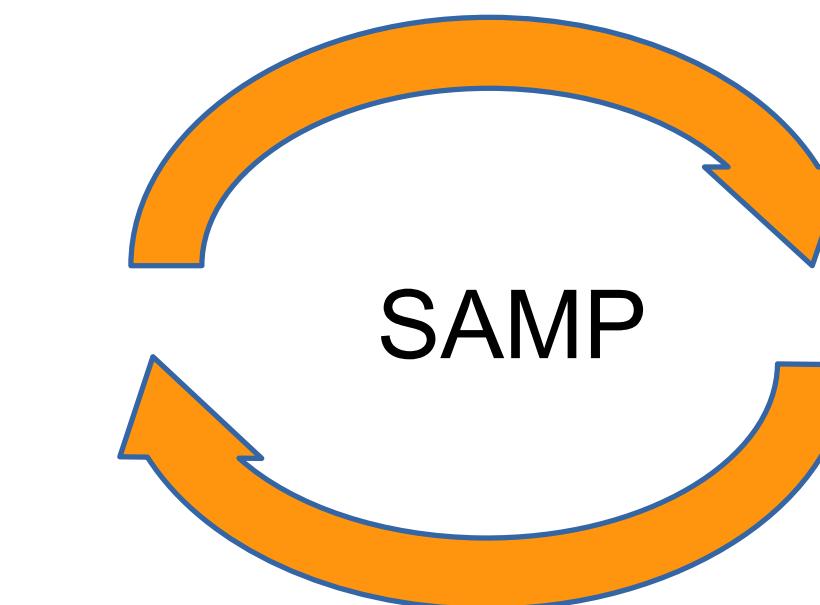
Aspro2

A screenshot of the Aspro2 software interface. It shows a map of the sky with a target star and various observation parameters like wavelength range, integration time, and sampling periodicity.

a2p2

A screenshot of the a2p2 software interface, showing a list of instrument configurations and a preview of a spectrum.

AMHRA

A screenshot of the AMHRA software interface. It shows a web-based analysis tool for high angular resolution, with sections for real-time astrophysical models and precalculated grids of astrophysical models.

Reduce  
data

amdlib  
pndrs

View Data

A screenshot of the OIFits Explorer software interface, showing plots of visibility data over spatial frequency.

OIFits Explorer

Fit Models

A screenshot of the LITPro software interface, showing a target panel and a fitter setup panel for model fitting.

LITPro

OiDB

A screenshot of the OiDB portal interface, showing a search form for astronomical objects and a results table.

L0 to L3  
DataBases

Results

A screenshot of the OiDB results table, showing a list of observations from 19 oifits files.

Reconstruct Images

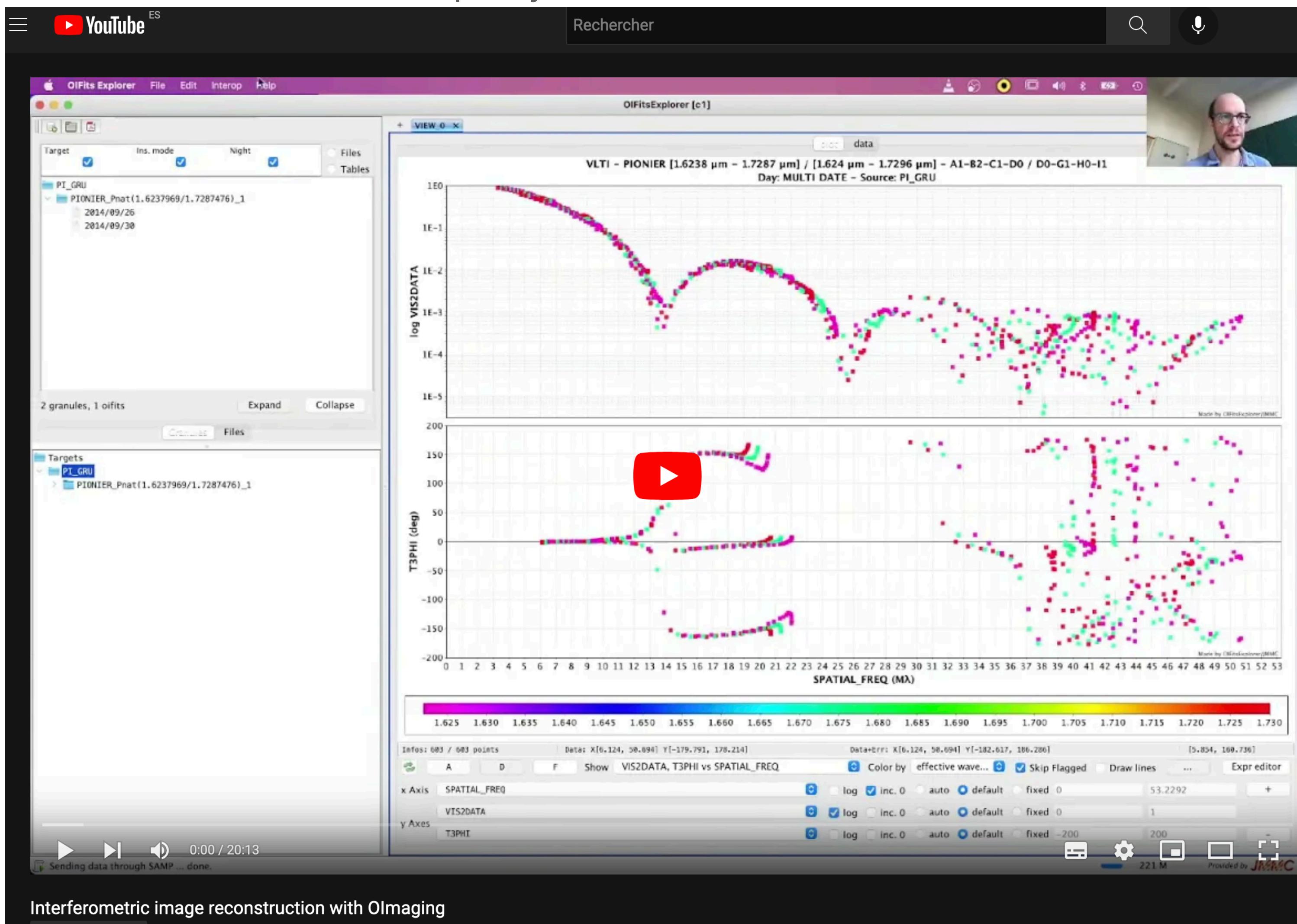
A screenshot of the Olmaging software interface, showing a reconstructed astronomical image and various processing parameters.

Olmaging

# Image reconstruction workflow

JMMC

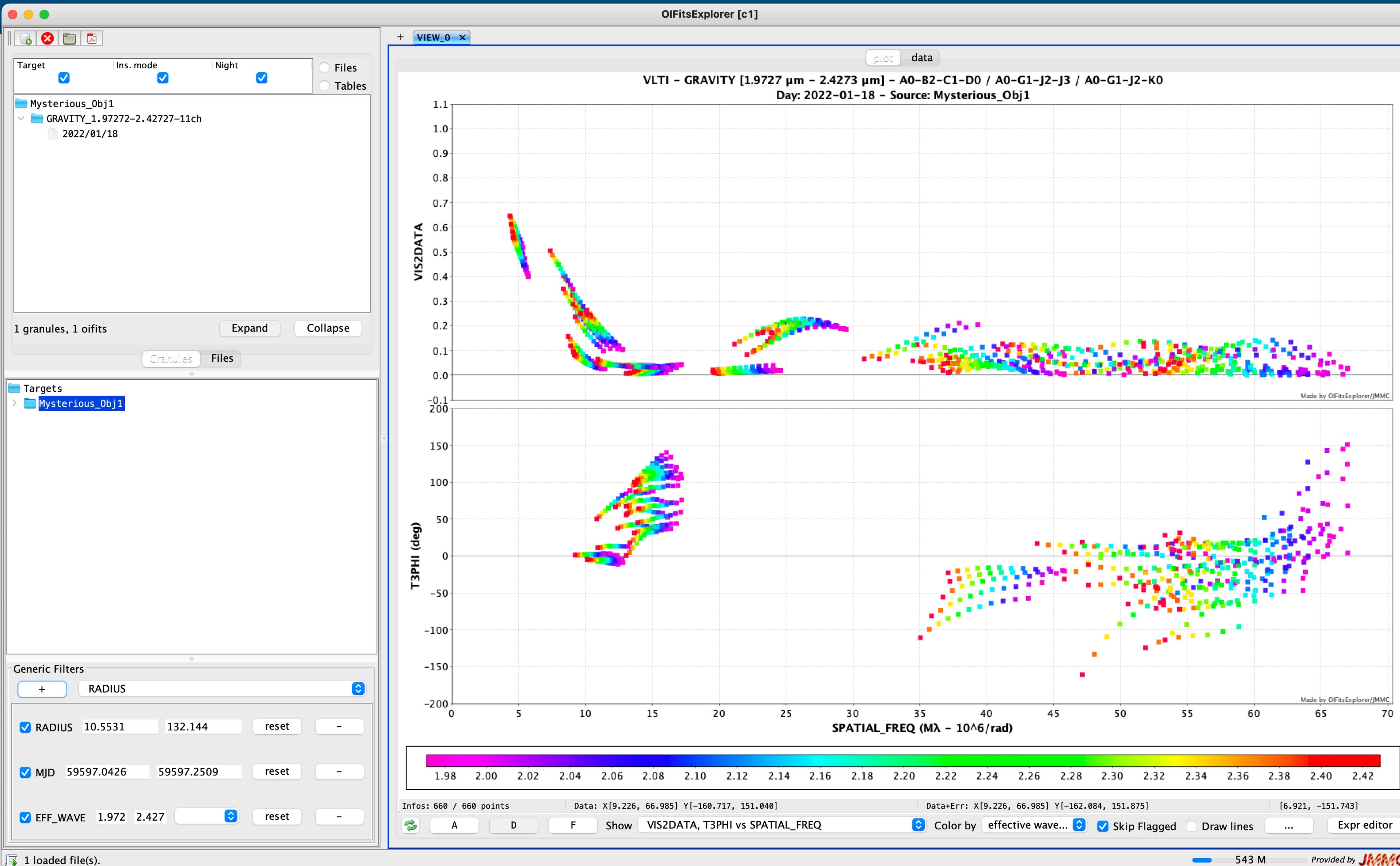
[https://youtu.be/YA3hvs\\_sOfE](https://youtu.be/YA3hvs_sOfE)



# OIFitsExplorer: data handling tool

JMMC

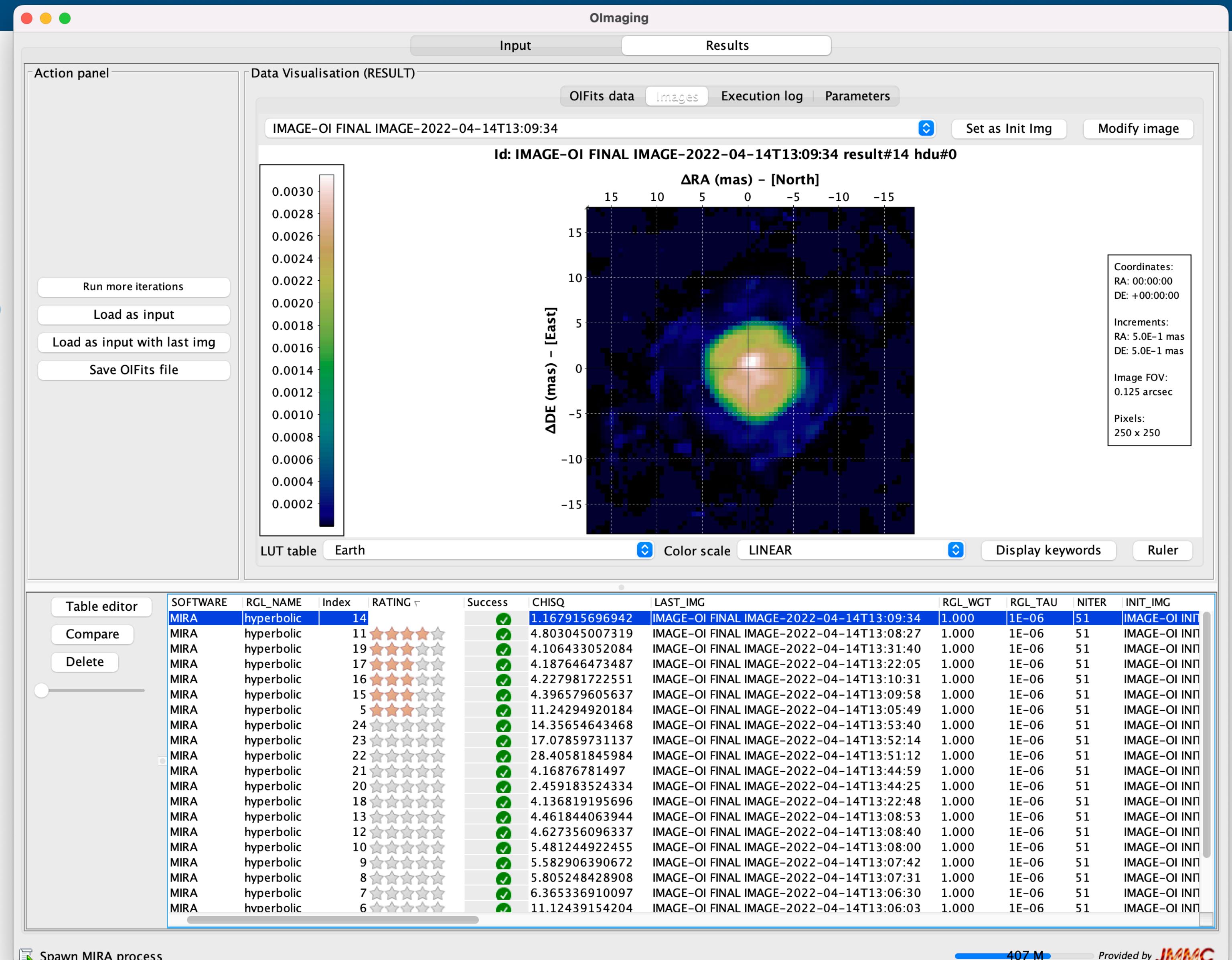
New version today!



## ◎ A single interface

- 4 softwares: BSMEM, MiRA, SPARCO, WISARD
- results in a single table
- rating, comparing,...
- saving reconstruction

parameters with the image

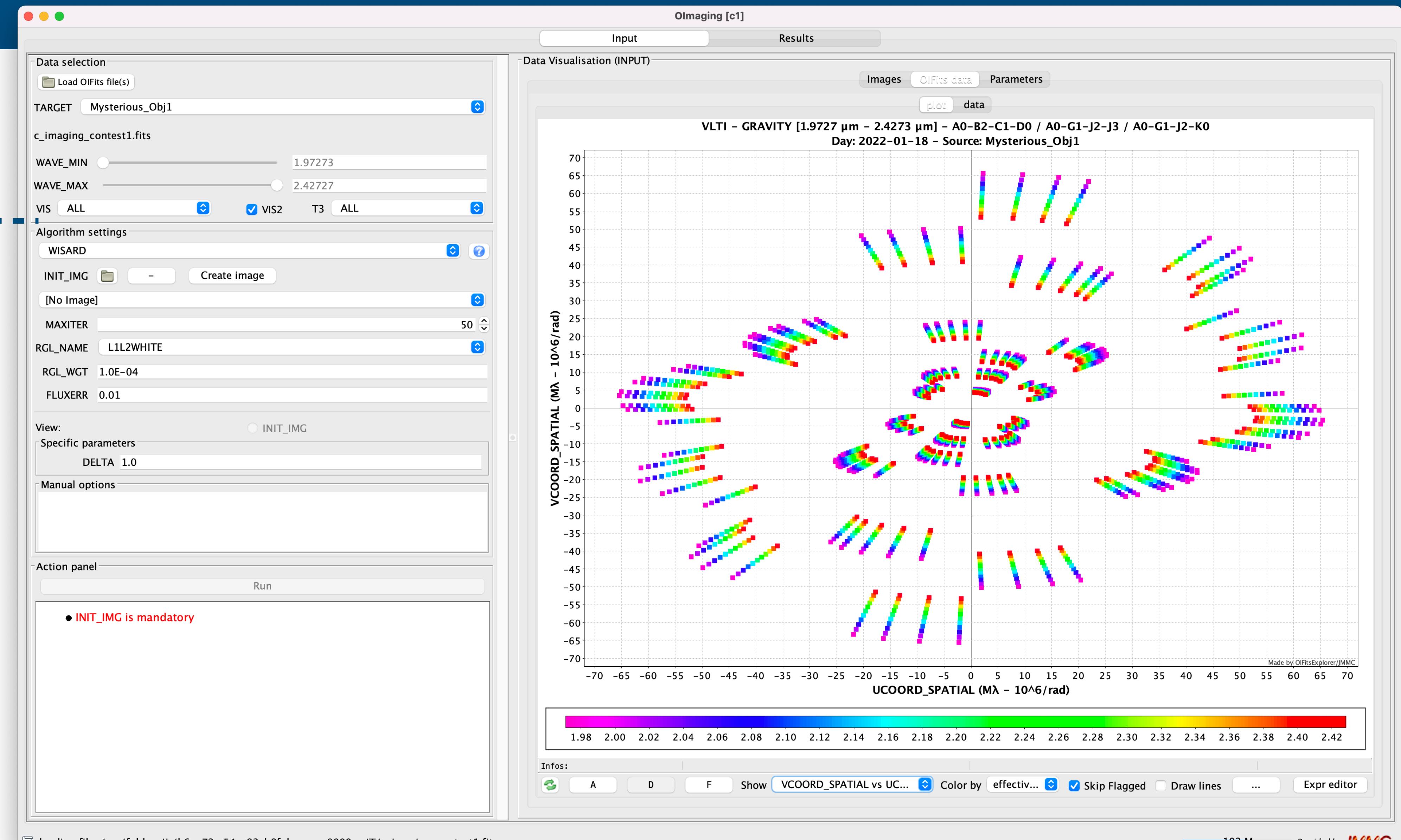


# Olmaging: Input panel

JMMC

## Data selection

## Algorithm parameters



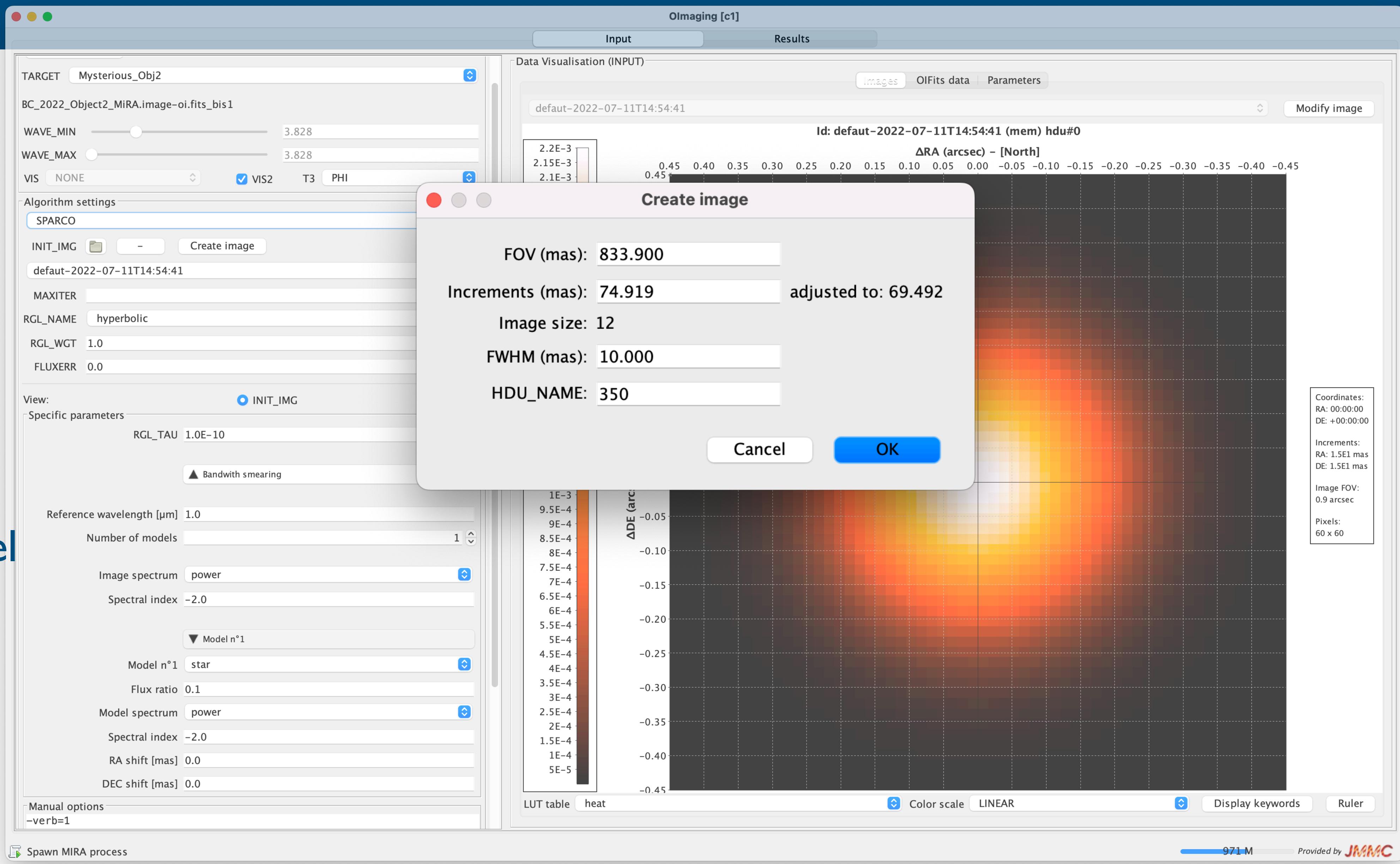
# Olmaging: Initial image

JMMC

Set the field of view and the sampling of the reconstruction

Can be a

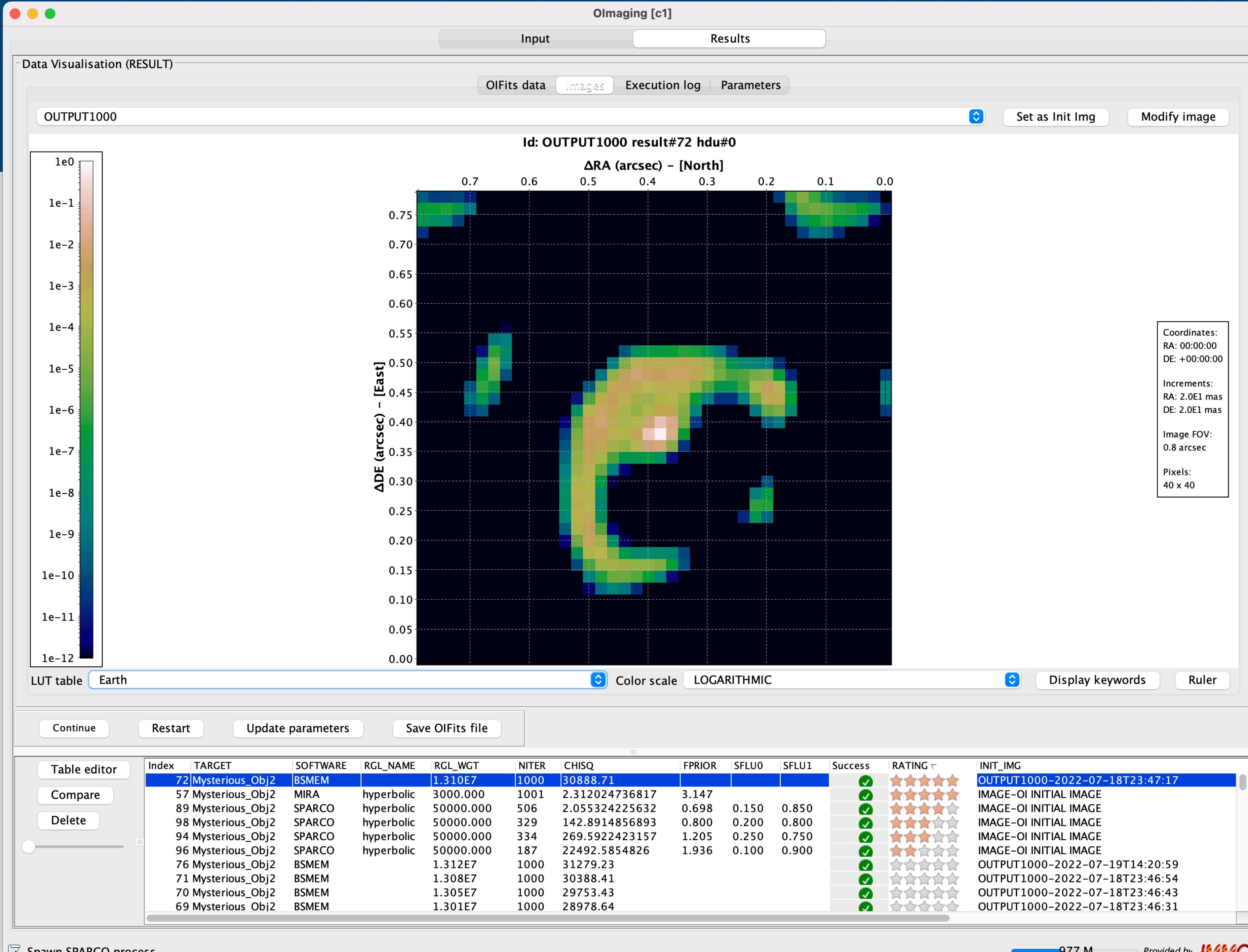
- FITS image
- image from model fitting algorithms
- built from Olmaging (Gaussian)



# Olmaging

## Result panel

- Image
- model
- execution log
- parameters



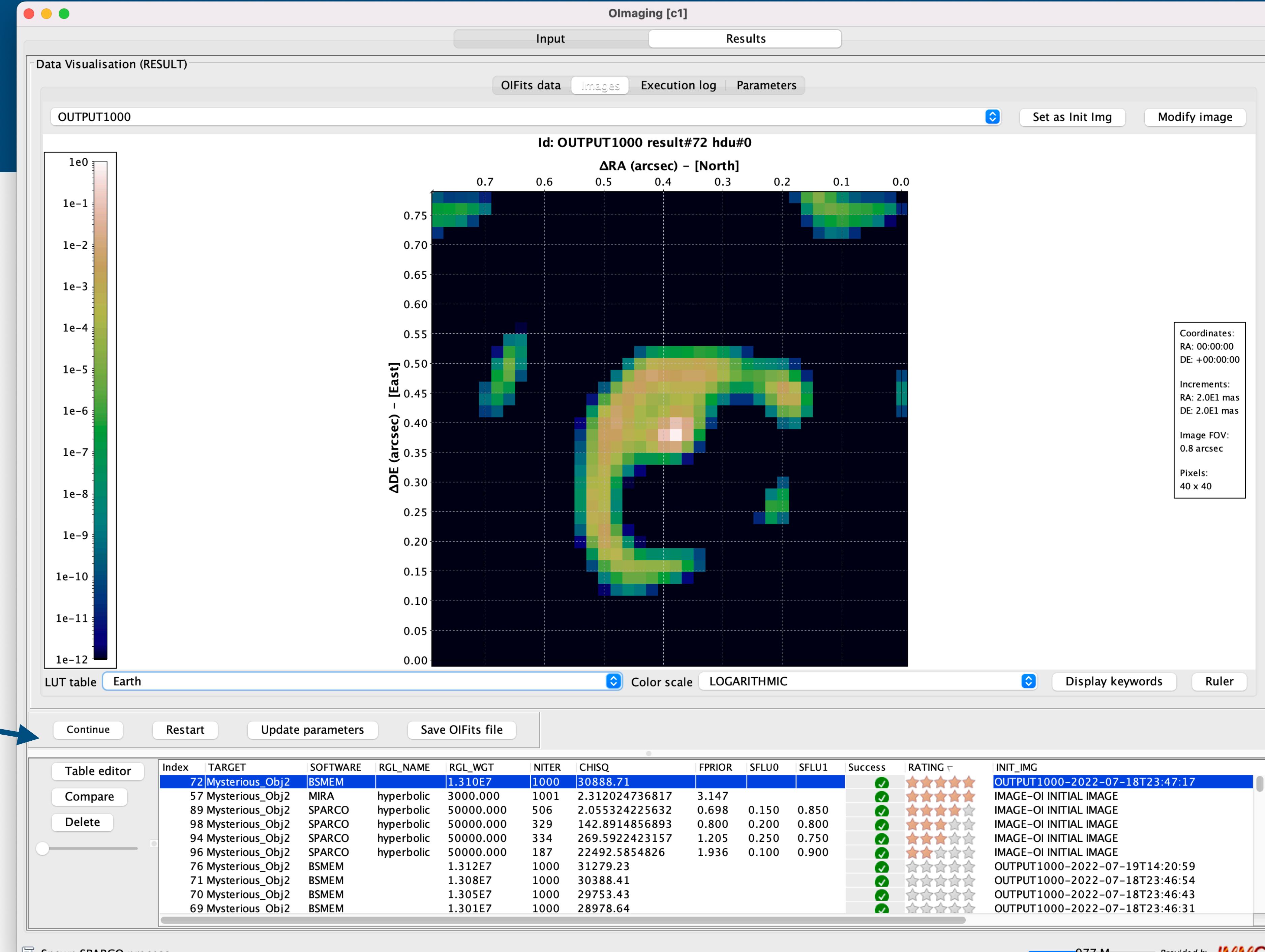
# Olmaging

## Result panel

- Image
- model
- execution log
- parameters

## Buttons

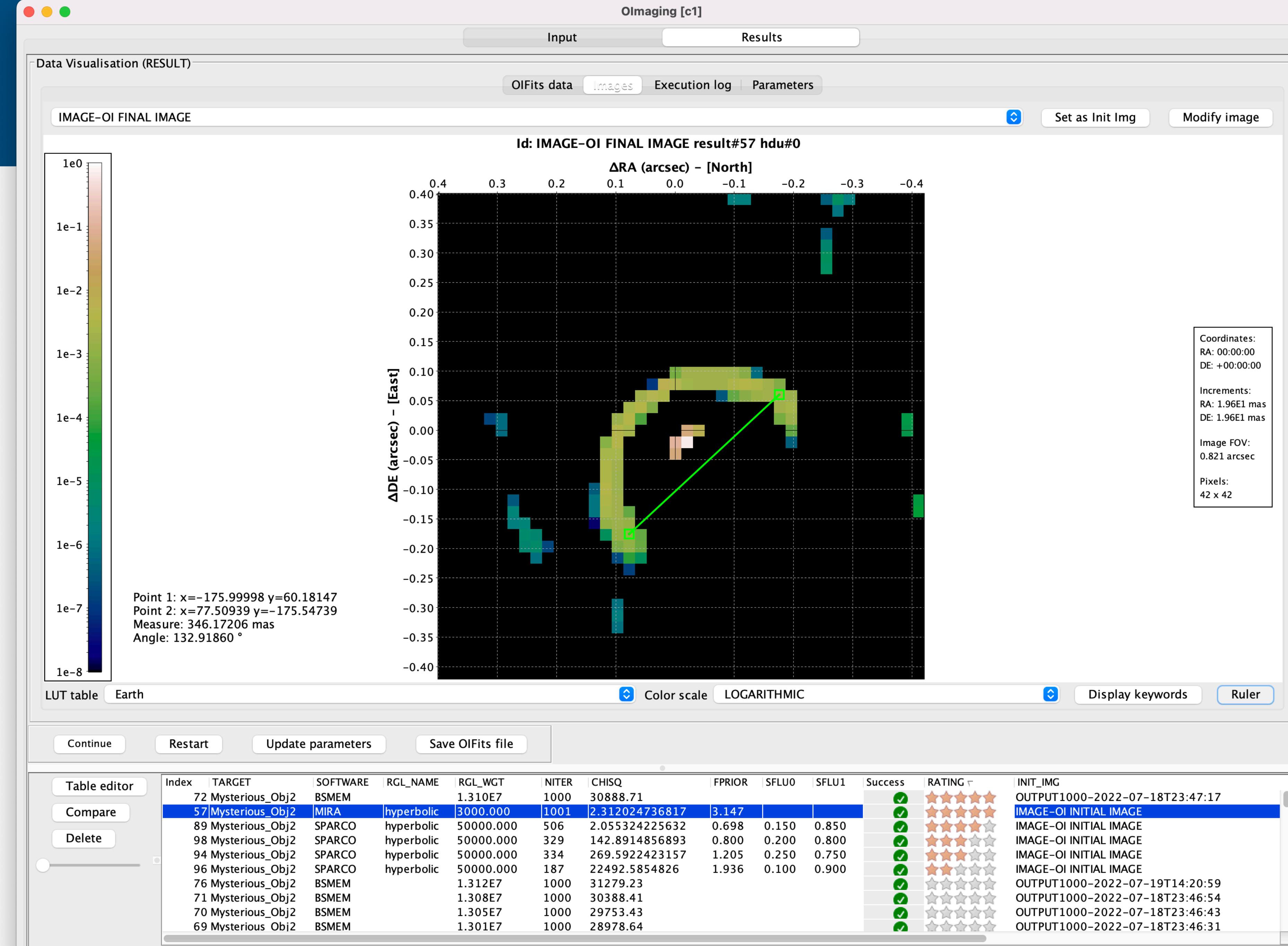
- Continue
- Restart
- Update
- Save



# Olmaging

## Tools

- ruler
- LUT table and scale



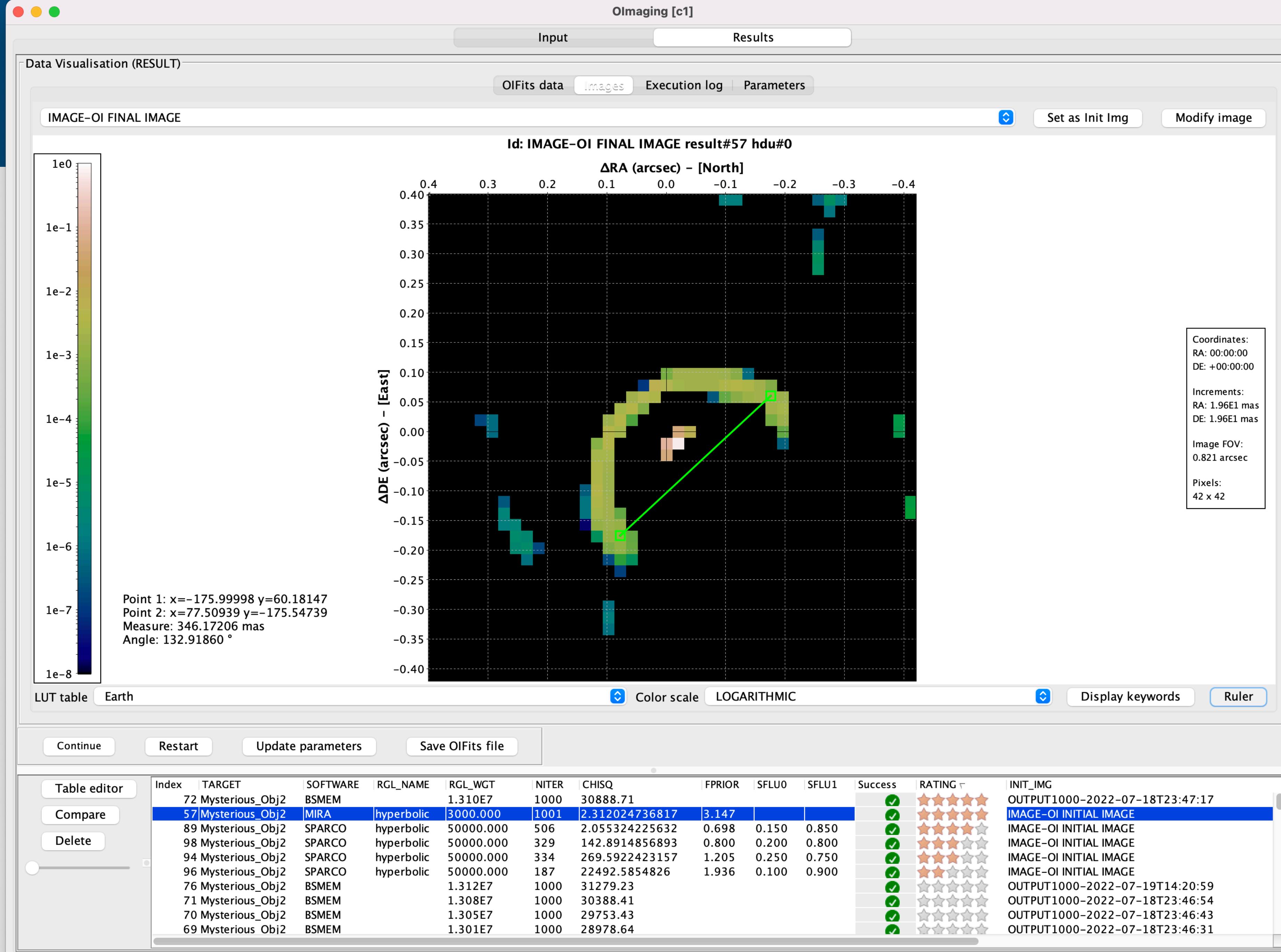
# Olmaging

## Tools

- ruler
- LUT table and scale

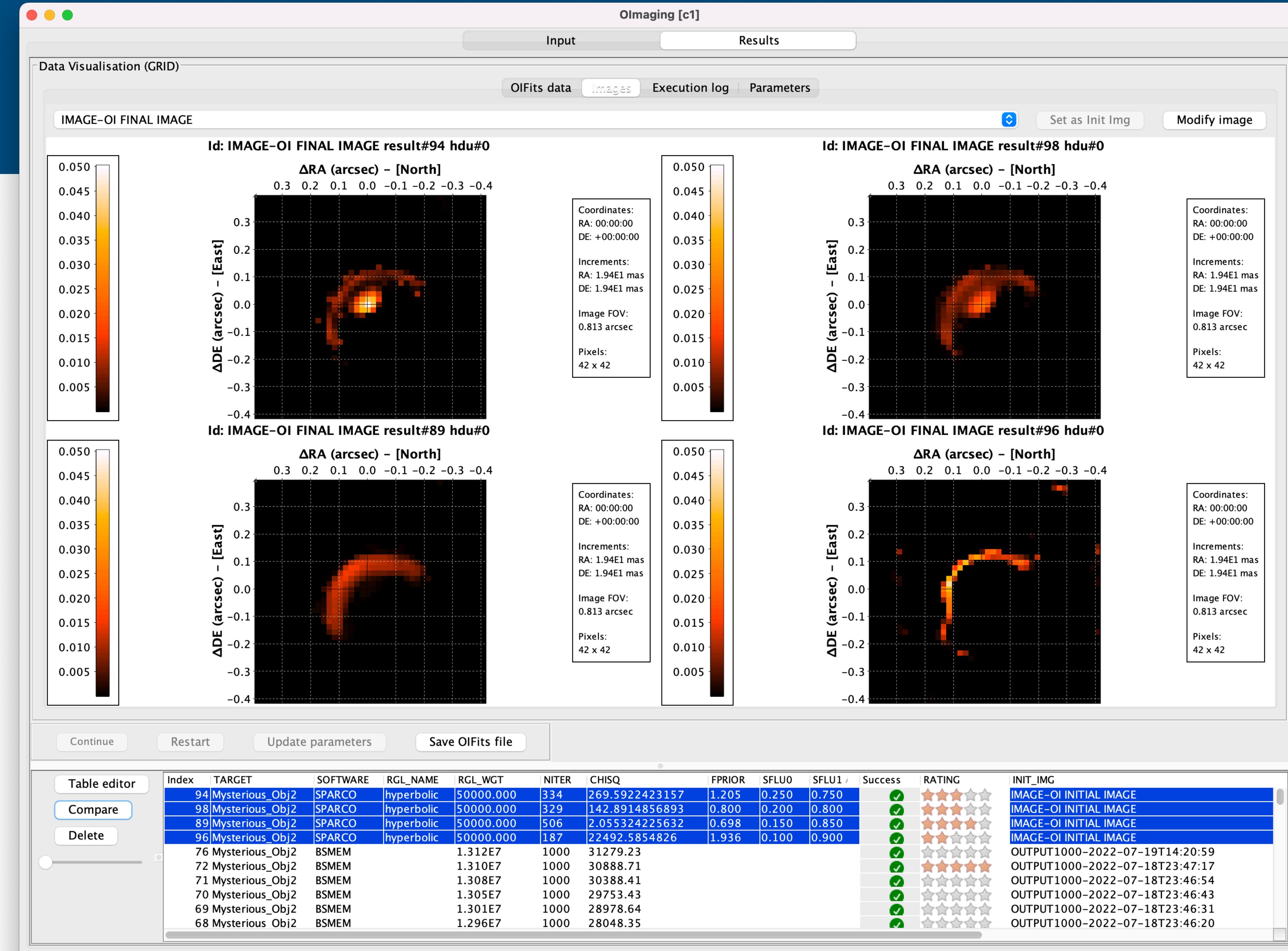
## Table panel

- All previous results
- sorting
- rating
- user notes



# Olmaging

## Comparing results



# You'll never walk alone

JMMC

Eii European Interferometry Initiative

Past Activities

VLTI Expertise Centres

Dissemination & Training

Joint Activities

Fizeau Program

Getting started with the VLTI

Subpages:

## VLTI Expertise Centres

Overview VLTI Expertise Centre Support

Structured development of optical interferometry requires leaping towards a European network of VLTI Expertise Centres. These centres are the backbone of dissemination activities to new VLTI users, by organising observing preparation and [data reduction schools](#), by co-organising with ESO VLTI open days, and being the end-points of the [Fizeau staff exchange programme](#).

The leap aims at bringing the impact and return of the programme in spreading know-how in Europe to a new level. It follows at a smaller scale the successful experience of the ALMA Regional Centres, where researchers travel to the expertise centres to reduce their data. The centres will be the visible first contact point for astronomers interested in using VLTI.

The present network of VLTI Expertise Centres includes three partners from the OPTICON Horizon 2020 networking activity:

- [Jean-Marie Mariotti Centre \(JMMC\) - Service aux Utilisateurs du VLTI](#), (SUV) France - a structure that aggregates manpower from different observatories:
  - [Observatoire des Sciences de l'Univers de Grenoble](#) (OSUG)
  - [Observatoire des Sciences de l'Univers de Lyon](#) (OSUL)
  - [Observatoire de Paris-Meudon](#) (OPM)
  - [Observatoire de la Côte d'Azur](#) (OCA)
- [Portuguese VLTI Expertise Centre](#), Portugal
- [University of Exeter](#), United Kingdom

two interferometry JRA (Joint Research Activities; WP8) lead partners:

- [Lagrange Laboratory/OCA](#), France
- [KU Leuven](#), Belgium

and two new nodes from the [OPTICON/RadioNet Pilot](#) (ORP) program:

- [Leiden Observatory](#), The Netherlands
- [Konkoly Observatory](#), Hungary

An overview of the support provided by each VLTI Expertise Centre and the data protection policy can be found [here](#).

Visitors wishing to travel to the above centres to reduce their VLTI data or prepare observations are encouraged to use the [Fizeau Programme](#).

### Welcome onto the JMMC User Feedback Form !

( \* : required field )

Application:

SUV (VLTI center)

Type:

Support Needed

Your Email \* :

your@email



Summary \* :

Comments \* :

Version:

Optional V.

[Effacer](#)

[Envoyer](#)

- Checkout new v1.0 release !!! [www.jmmc.fr/releases/](http://www.jmmc.fr/releases/)
- Still in progress
- Open source:  [github.com/JMMC-OpenDev](https://github.com/JMMC-OpenDev)
- Need help to add new algorithms

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- Need help to add new algorithms

## Ask for help and give feedback !

- Expertise center gather experts that are willing to help
- Request features
- Knowing user needs help us to tailor softwares
- Research on methodology are fueled by requests

# Olmaging: a collective project

JMMC

## ○ The big chiefs:

I. Tallon-Bosc  
J-P. Berger  
G. Duvert

## ○ The developers:

L. Bourgès  
A. Kaszczyc  
G. Mella  
M. Pratoussy



## ○ The reconstruction software fathers:

G. Duvert  
J. Kluska  
L. Mugnier  
E. Thiébaut  
J. Young

## ○ The beta-testers:

J. Kluska  
M. Montargès