

SOFIA – 2.5m IR telescope at 13,000m

Herschel – 3.5m FIR telescope at 1.5 m Km



First light
2009!



Future Technologies for Telescopes and Instruments

Colin Cunningham

Director, UK Extremely Large Telescope Programme



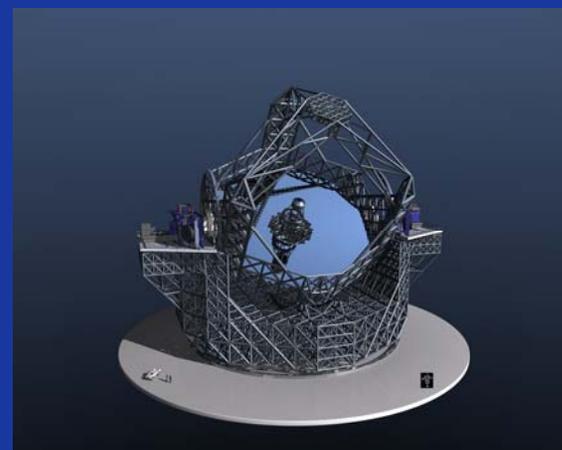
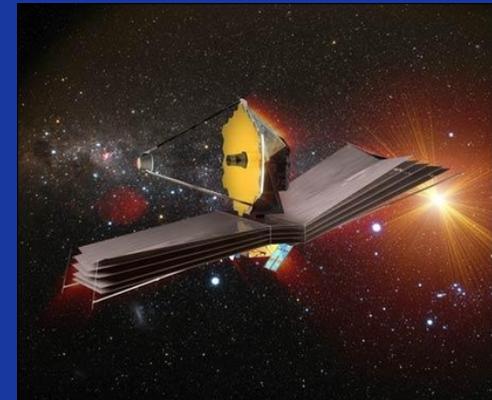
Science & Technology Facilities Council
UK Astronomy Technology Centre

Royal Observatory Edinburgh



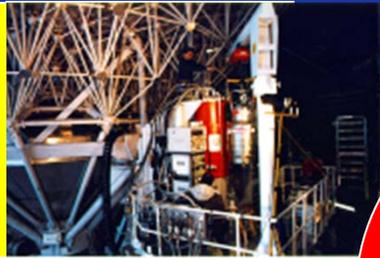
Scope of talk

- Innovation in Optical/IR Astronomy – ground & space
 - Science Drive and Technology Push
 - Disruptive Technologies
 - Gestation Periods for new technology to be adopted
- Survey of current ideas for new technology for telescopes and instruments
- Where might we go in next 50 years?



Why do we need Technology Development in Astronomy?

**Open New
Parameter Space
> New Science**



**Disruptive Technology
– changes how we do things**

**Reduce cost- make
New Science affordable**



**Reduce time to
reach Science Goal**



Reduce Risk

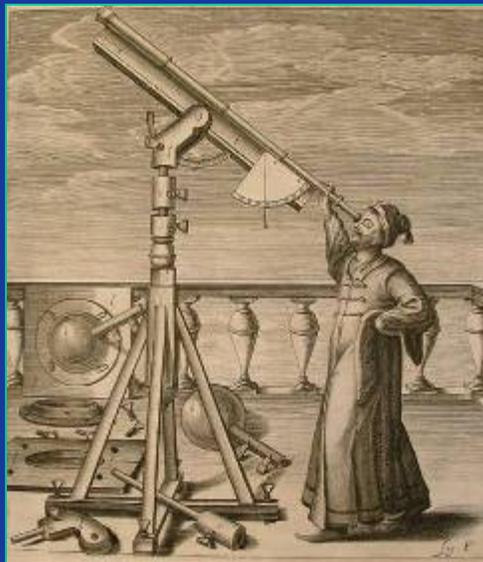
Disruptive Technology

A new technology that has a serious impact on the status quo and changes the way people have been dealing with something, perhaps for decades



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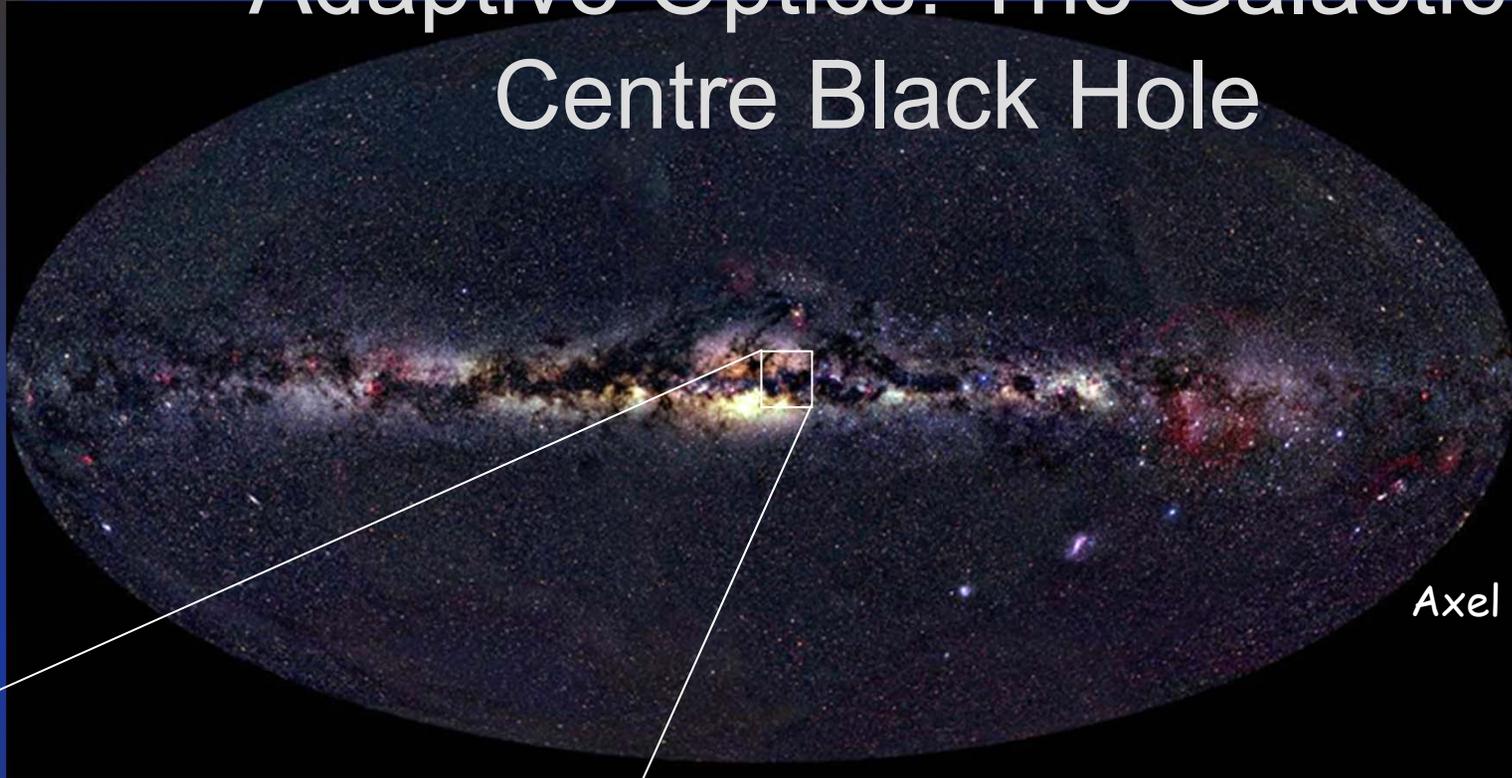
Credit: LSST Corp. 5

Which Technologies have changed how we do astronomy?

- Lens > Galileo's Telescope
- Metal mirrors > Herschel's Telescope
- Silver on glass mirrors > Foucault's Telescope
- Go up a mountain > Tenerife, Lick Obs
- Pyrex Glass Mirror & Ship engineering > Hale 200" Telescope
- Electronic Detectors: CCDs & IR Arrays
- Space Telescopes > HST
- Segmented Mirrors > Keck
- Active optics > VLT, Gemini Subaru



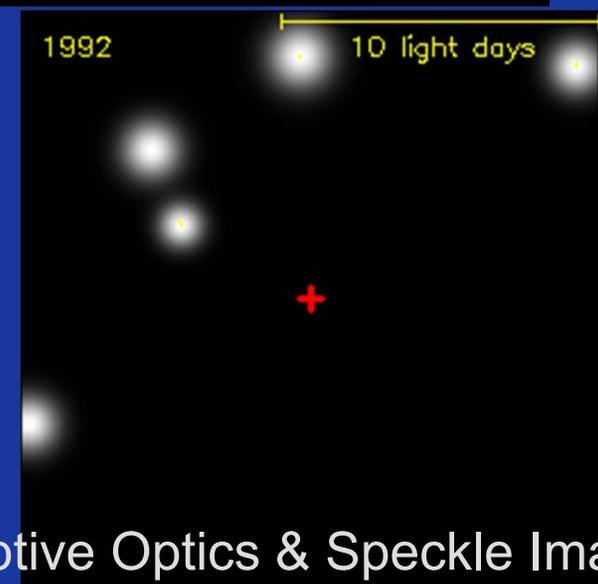
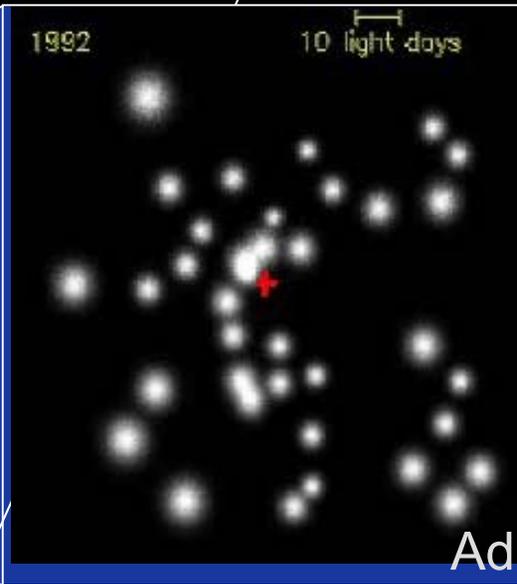
Adaptive Optics: The Galactic Centre Black Hole



Axel Mellinger



ESO VLT
Genzel et al., 2003



Adaptive Optics & Speckle Imaging

What will the Disruptive Technologies of the future be?



By definition – impossible to answer!

**"If you can look into the
seeds of time, and say
which grain will grow and
which will not, speak then
unto me"**

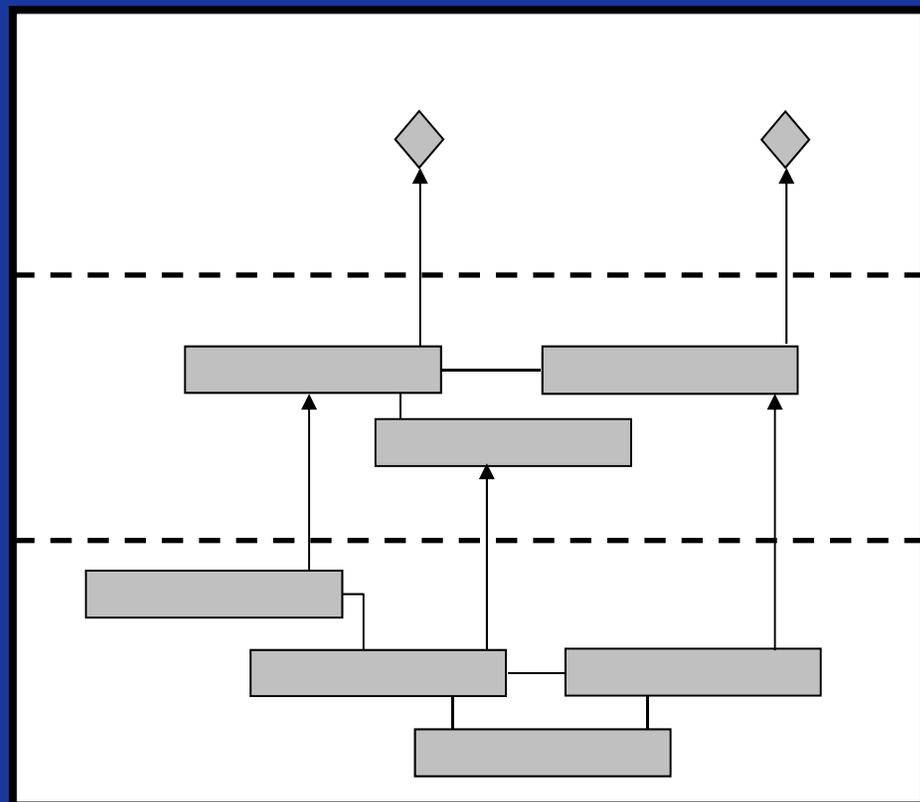
*William Shakespeare
(Macbeth)*

Technology Planning: Roadmap

Science Goals

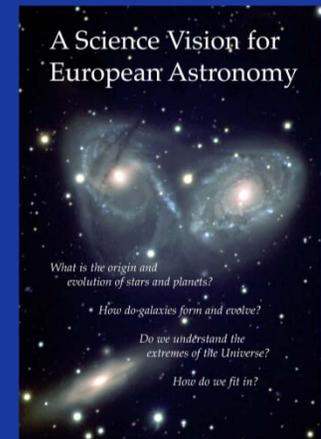
Facility, Mission or Instrument

Technology



Science Drivers

- NASA 2006 Strategic Plan
 - Discover the origin, structure, evolution, and destiny of the universe, and search for Earth-like planets
- ESA 2008 Cosmic Vision
 - What are the conditions for life and planetary formation?
 - How did the Universe begin and what is it made of?
- ASTRONET 2008 European Science Vision
 - The Extremes of the Universe
 - What is Dark Matter & Dark Energy?
 - Galaxy Formation & Evolution
 - How were galaxies assembled?
 - How did our Galaxy form?
 - How do galaxies form and evolve?
 - Origins and evolution of stars and planets
 - Where are most of the metals throughout cosmic time?
 - How do we and the solar system fit in?



Science Needs Drive Technology Requirements

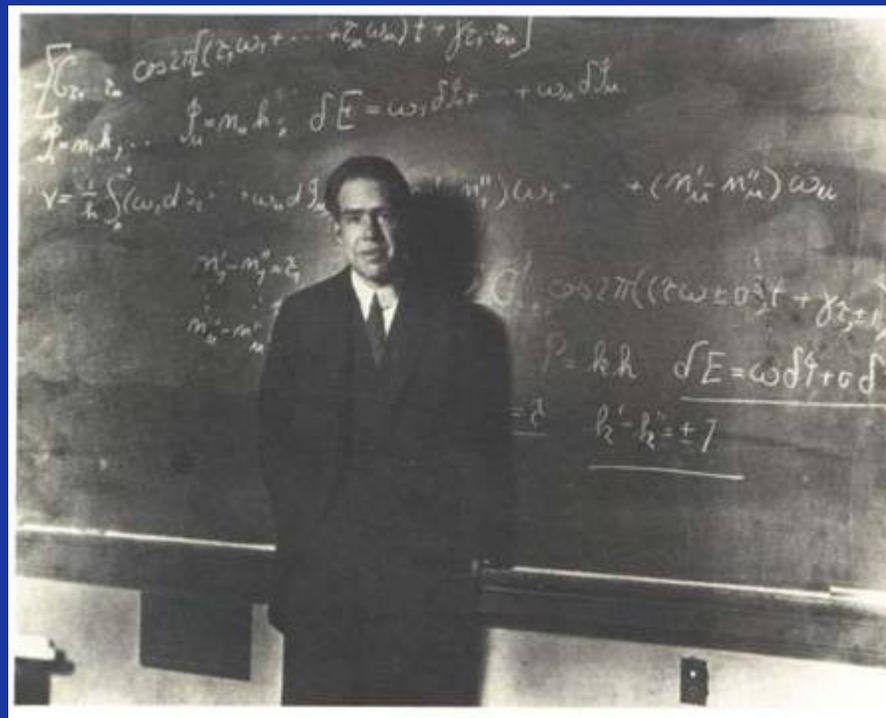
- Is it is simple as this?
- Can we define science needs, then facility requirements, then decide what technology should be developed and when?
- Probably yes, for evolutionary developments
- Or is it often much more iterative?
 - Disruptive technologies push us into new parameter space or observing methods
 - Detectors
 - Computing
 - Photonics?

Political Drivers

- Why do governments spend money on astronomy?
 - Gain or maintain World standing in science
 - Feed Industry
 - Inspire people into science and technology
- Organisational agenda: SIRTf/Spitzer & NASA
- We don't succeed without taking care of these issues

Predicting the next 50 years

"Prediction is very difficult, especially if it's about the future." Nils Bohr



Or was it Yogi Berra?

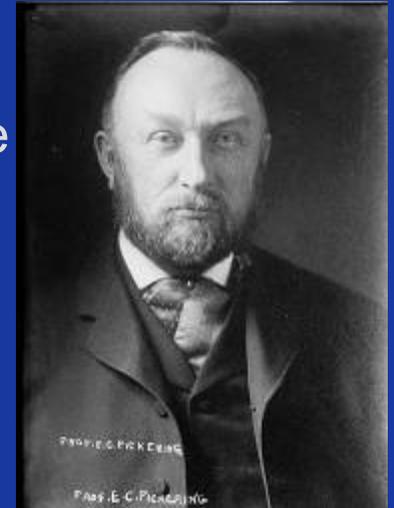
“It's tough to make predictions,
especially about the future”



Prediction from 100 yrs ago

Edward Pickering 1909 POPULAR
SCIENCE MONTHLY

- ‘Have we at length reached the limit in size?.... It is more than doubtful, however, whether a further increase in size is a great advantage. Much more depends on other conditions, especially those of climate, the kind of work to be done and, more than all, the man behind the gun’
- Other predictions:
- 1) Service Observing at 2 Large observatories at +30 deg and -30 deg, with central data reduction centre
- 2) Determining stellar distances
 - Differential absorption in interstellar medium
 - Variations in speed of light with wavelength !!!
 - ‘If we could determine our motion with reference to the ether, we should have a fixed line of reference to which all other motions could be referred’

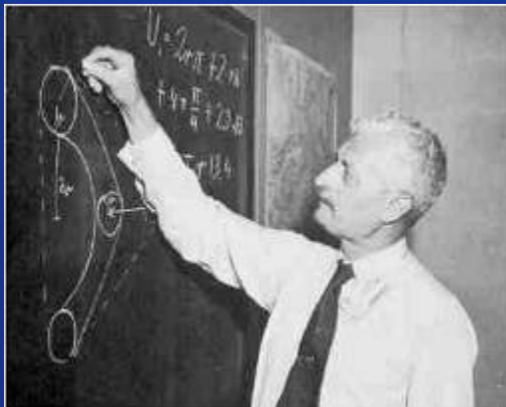


Prediction from 50 yrs ago

Henry King, 1955

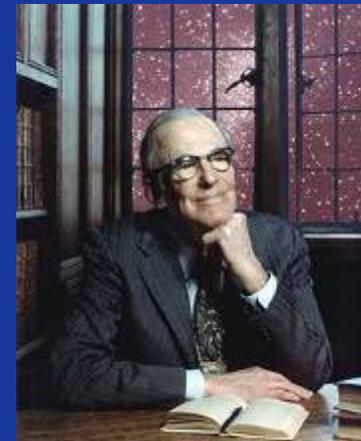
The History of the Telescope (pub Griffin)

'Thus we return to the old problem of escaping from the damaging optical effects of the earth's atmosphere – and there is no escape as yet....in the future, when interplanetary travel becomes possible, observing stations and, ultimately, observatories may be established in space'



Observatory in space
was proposed in
1923 by Hermann
Oberth

Lyman Spitzer
described benefits
of a space
telescope in 1946



What can we measure ?

I
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Why is this so hard?

Probe the Universe by collecting photons:

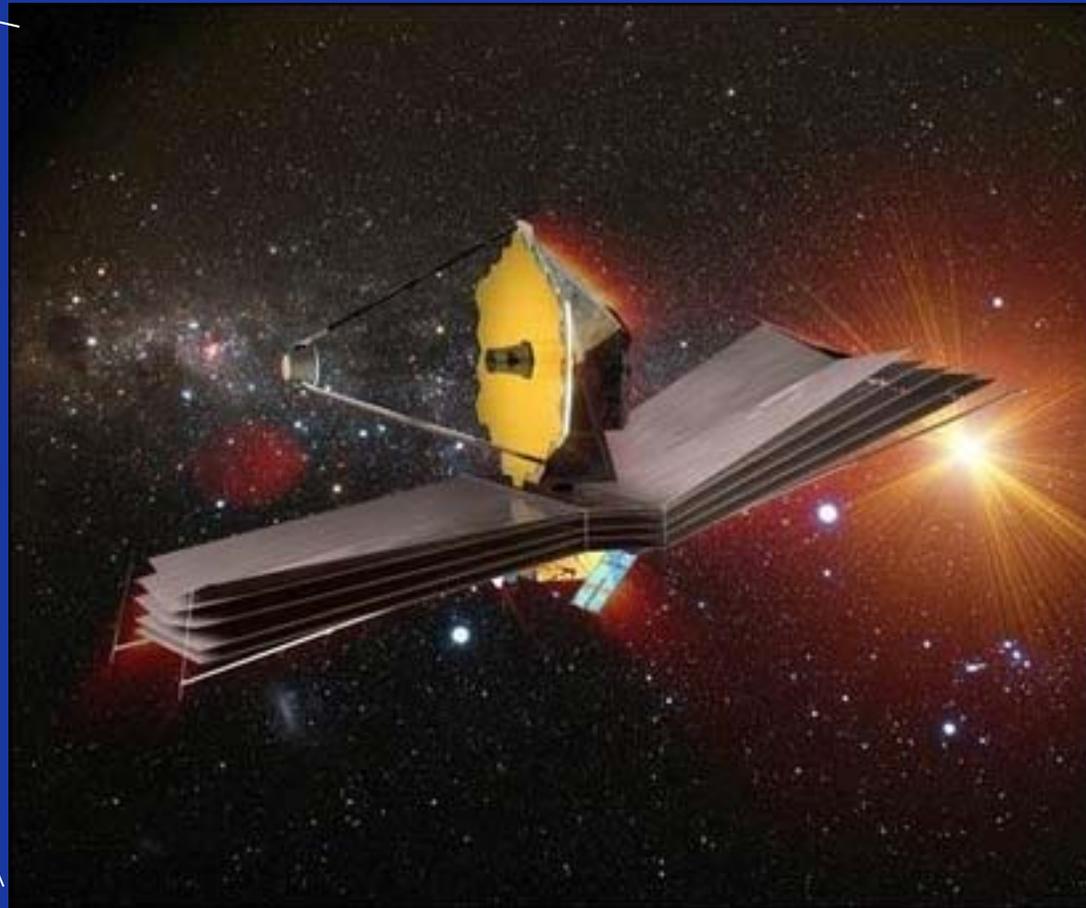
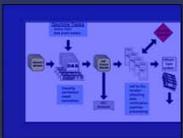
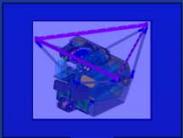
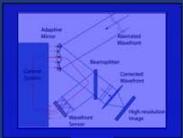
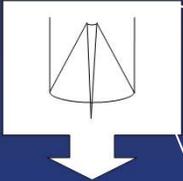
- Photon emission
- Absorption
- Eclipses
- Gravitational lensing
- Angles - astrometry
- Doppler shift > dynamics of galaxies > Dark Matter
- Measure properties of E-M radiation:
 - Intensity
 - Spatial
 - Temporal
 - Wavelength (Energy)
 - Amplitude & Phase
 - Polarisation
 - Orbital Angular Momentum?

We can also measure or collect:

- Neutrinos
- Gravitational Waves
- Cosmic rays (particles)
- Meteorites



Telescopes



What does the telescope structure do?

Example: Gemini Telescope

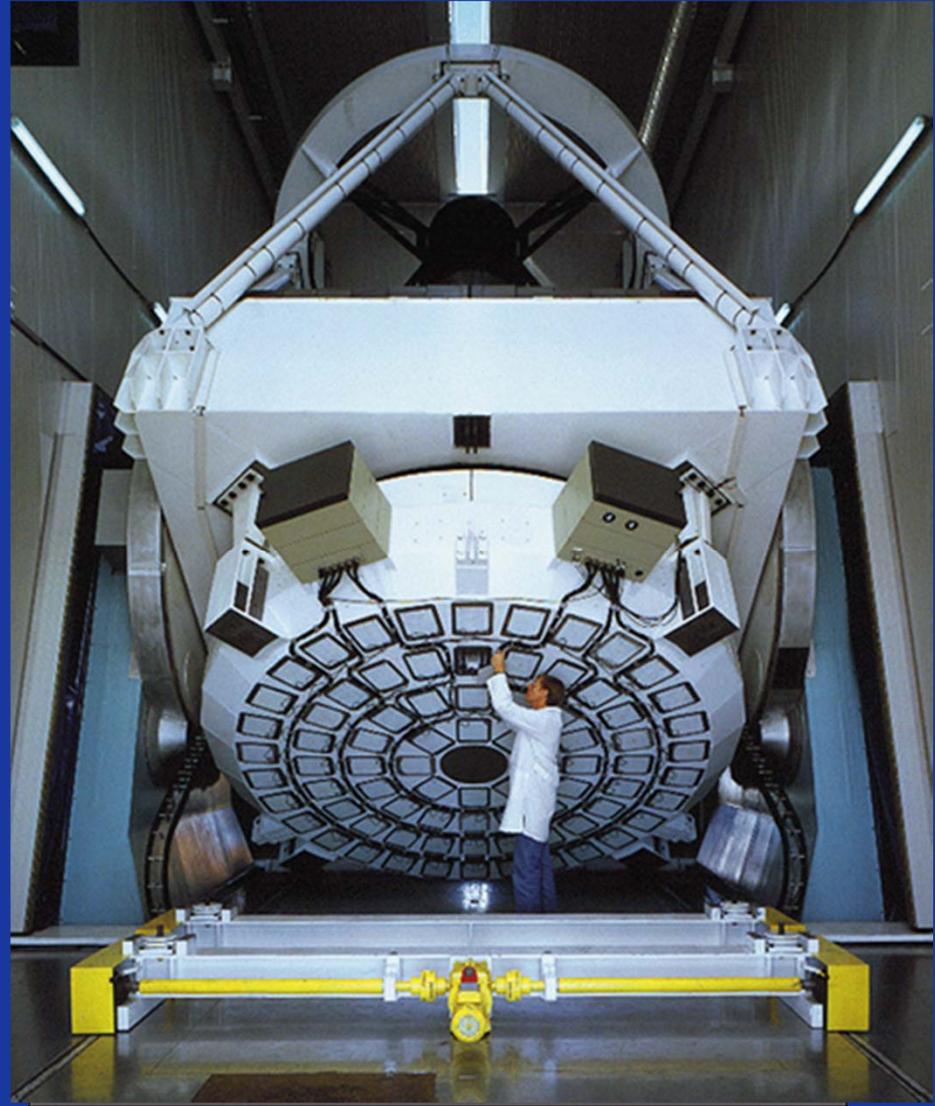
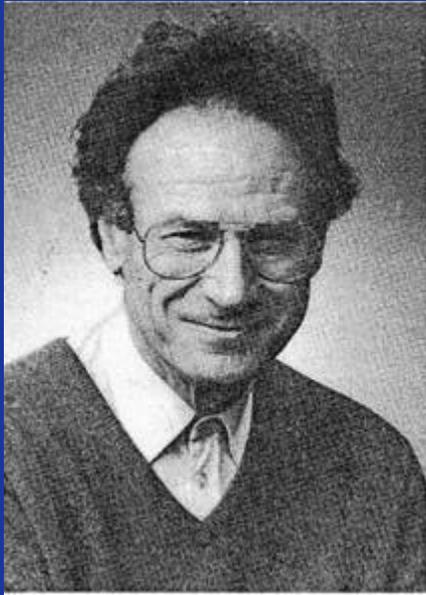
17g of aluminium or silver supported by 380 tons of structure!

How about making the structure 'smart' ?

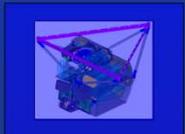
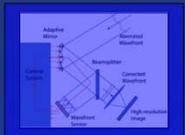
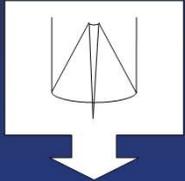


Gemini Observatory

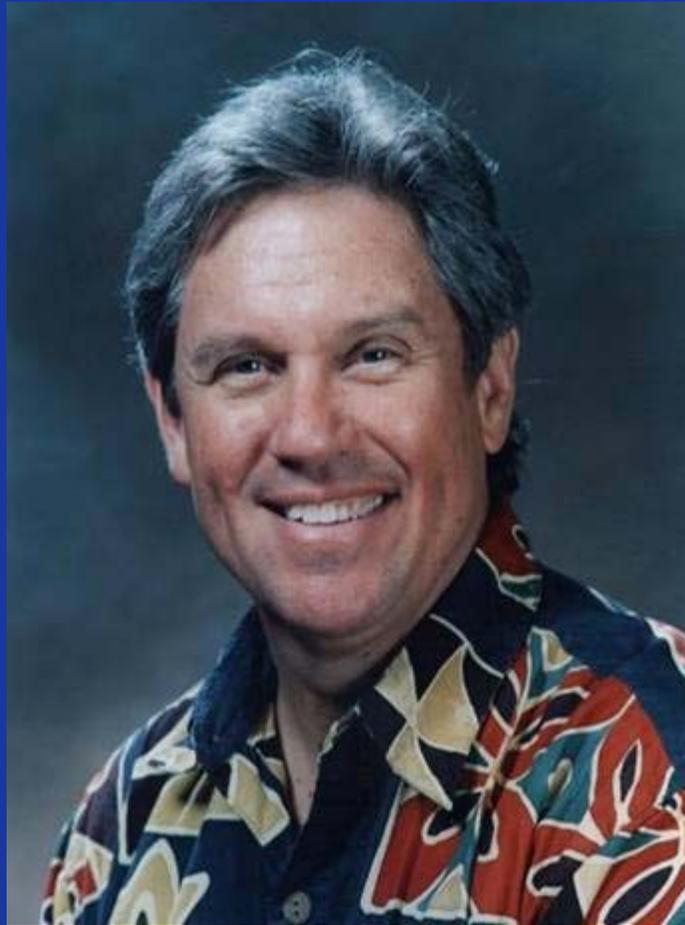
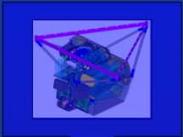
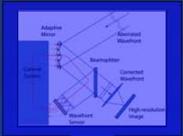
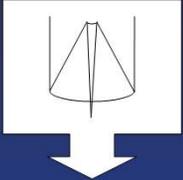
Ray Wilson & Active Optics



- Active Optics
 - Thin Mirror
 - Wavefront Sensing
 - Actuators
- New Technology Telescope>
 - VLT
 - Subaru
 - Gemini



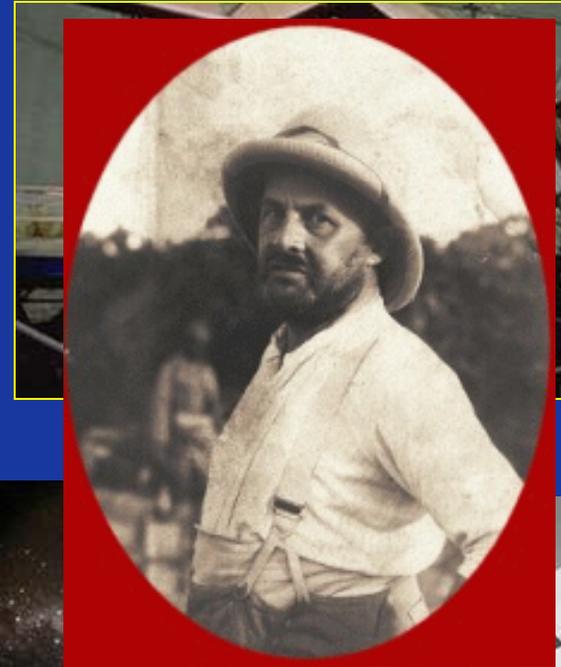
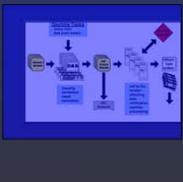
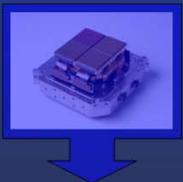
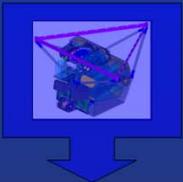
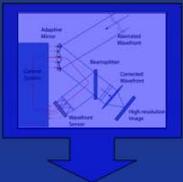
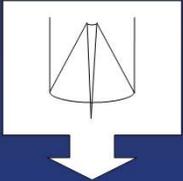
Jerry Nelson & Segmented Telescopes



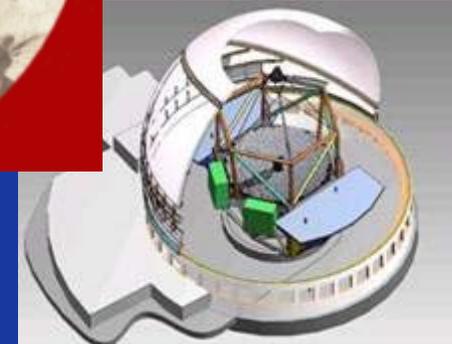
10m Keck Telescopes: 36 Segments

©Russ Underwood/W.M. Keck Observatory

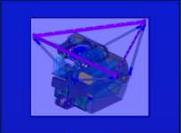
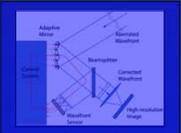
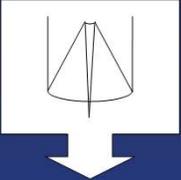
Segmented Mirrors



- Not new!
- Guido Horn d'Arturo at Bologna Observatory
- Built between 1935-52
- 61 segments to form 1.8m diameter mirror



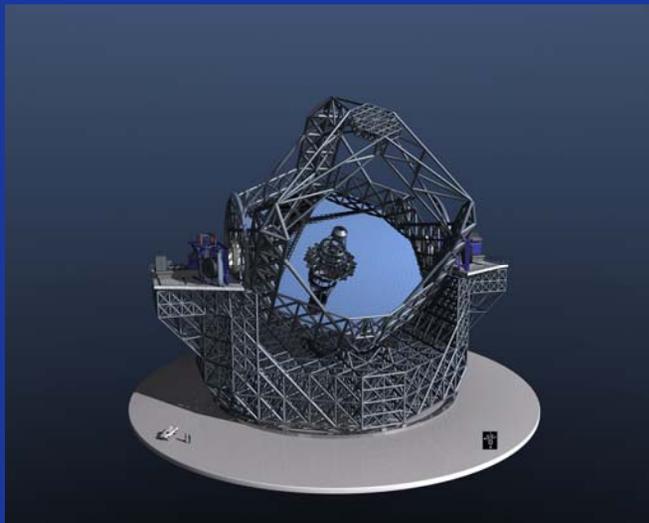
Enables next step: bigger telescopes!



Thirty-Meter Telescope Project

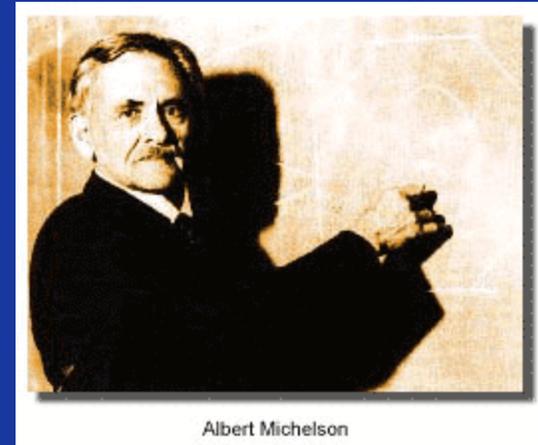
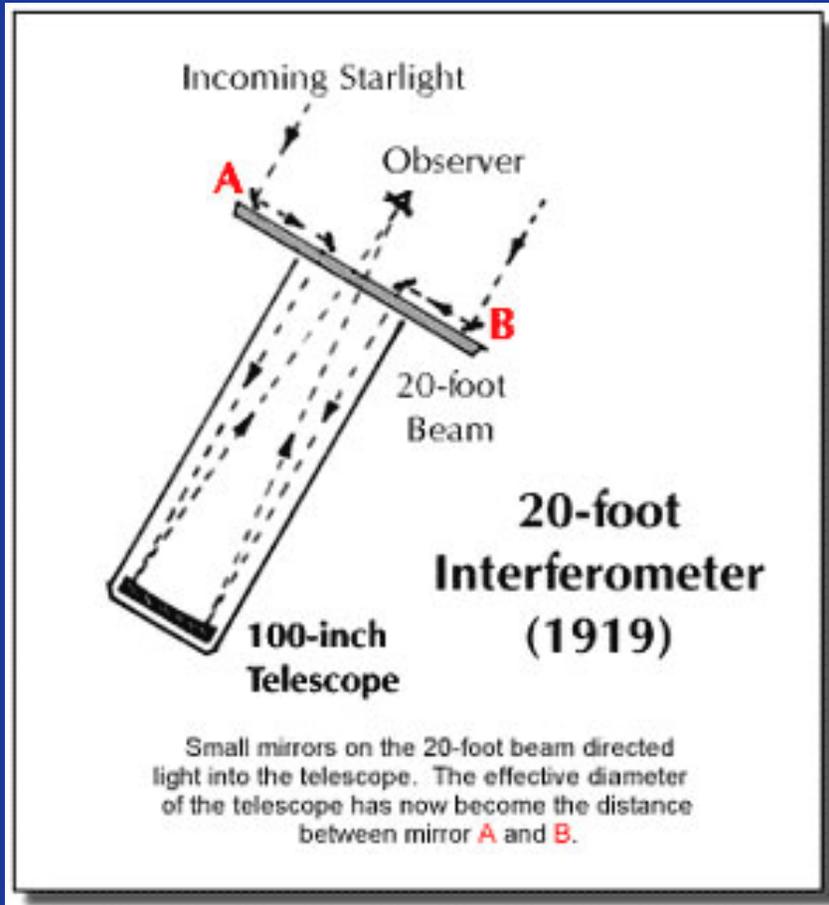
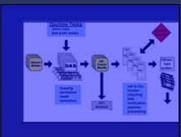
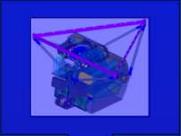
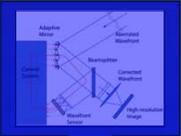


Giant Magellan Telescope - Carnegie Observatories.
Artwork by Todd Mason



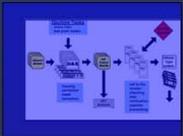
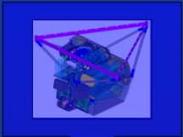
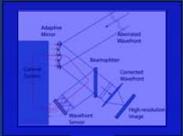
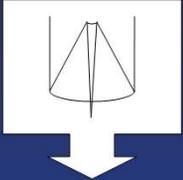
European Extremely Large Telescope

Increase apparent D: Interferometers



The 20-foot beam on top of the 100-inch Hooker Telescope on Mt. Wilson in Southern California.

Interferometers : Novel technologies



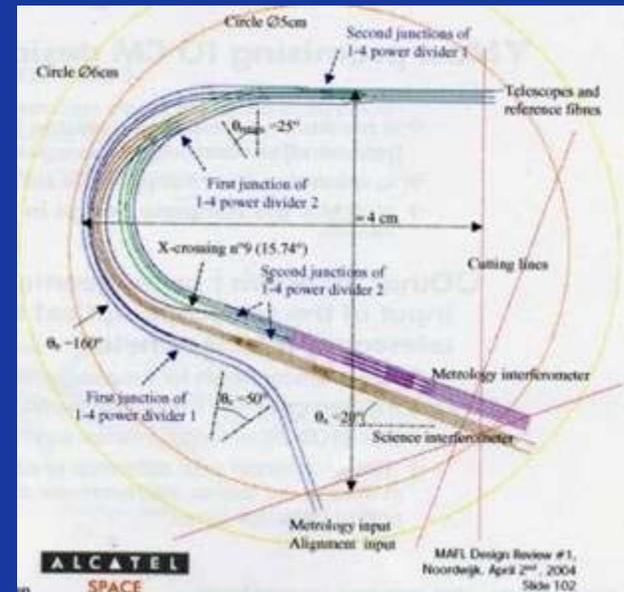
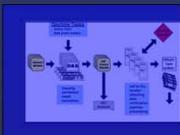
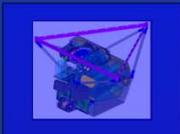
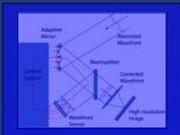
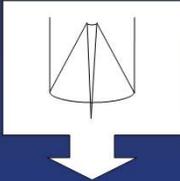
- Instruments are complex and large
- So integrated optics are attractive



Full integrated instrument: Multi Aperture Fibre Linked Project

An integrated instrument that includes most of the optical functions required to control a three-arm interferometer:

- ✓ coherent transport including modal filtering
- ✓ metrology system
- ✓ 3-channel beam combiner



SPACE: Sputnik Changed Everything!

This geopolitical earthquake fundamentally changed the way we look at science and science funding



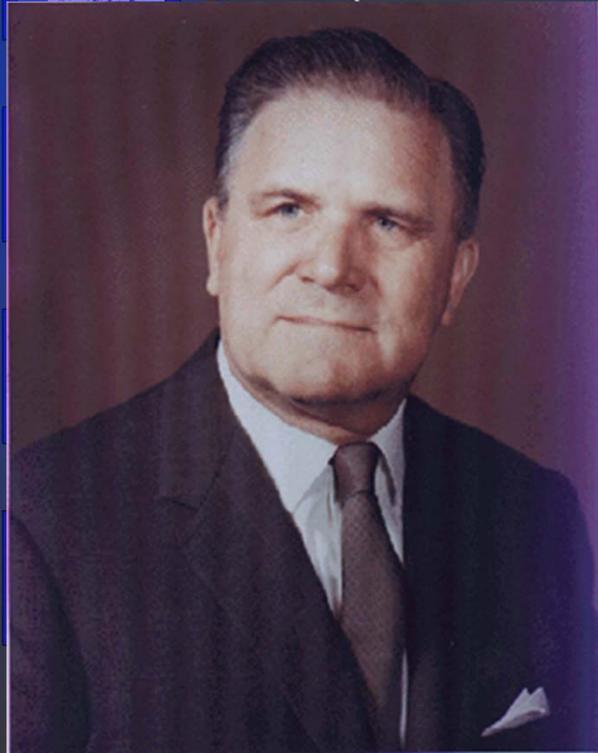
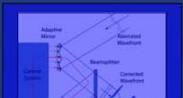
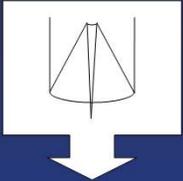
1957 NY Times front page

Jim Crocker,
Lockheed Martin



Launched 1990

Fixed 1993



Optical Telescope Element (OTE)

Integrated Science Instrument Module (ISIM)

ng side

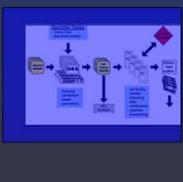
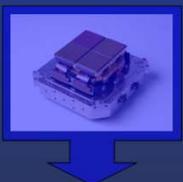
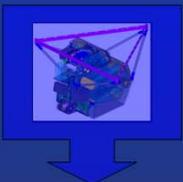
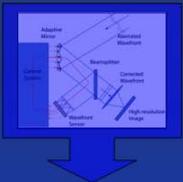
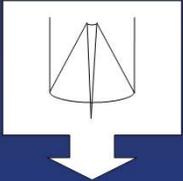
Warm, Sun-facing side

Spacecraft Bus

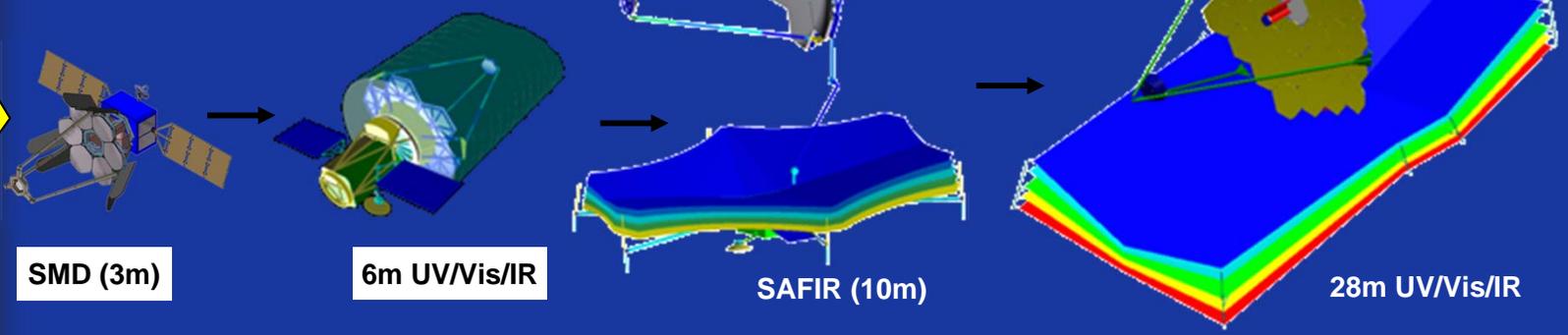
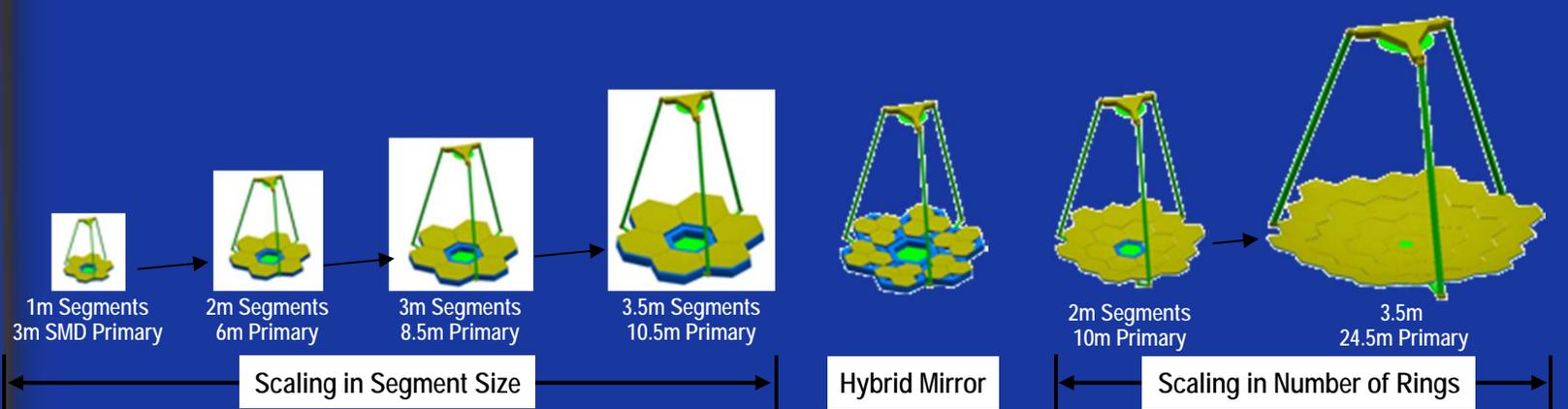
James Webb Space Telescope

Possible Next Steps for Space Telescopes

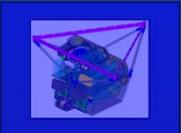
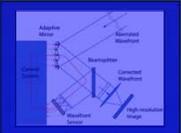
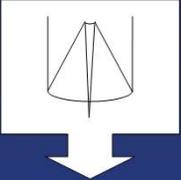
- Larger filled aperture telescopes
 - Using ARES V launcher
 - Could launch 8m monolithic mirror
- Robotic assembly of giant telescopes
- Nulling Interferometers
 - ESO: Darwin
 - NASA: Terrestrial Planet Finder



Scaling to Very Large Apertures



ARES V & 8m monolithic Telescope



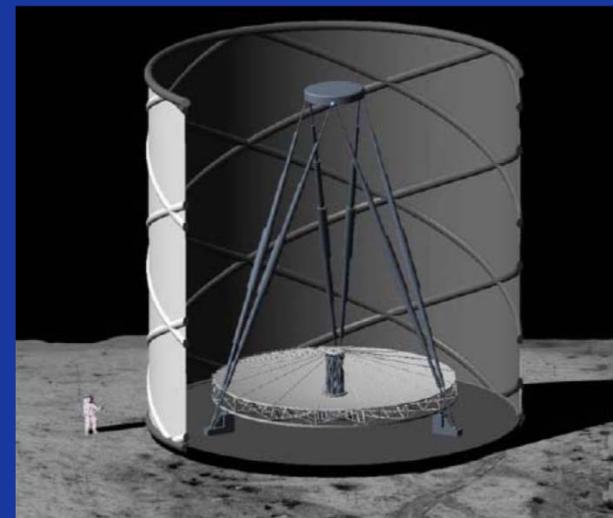
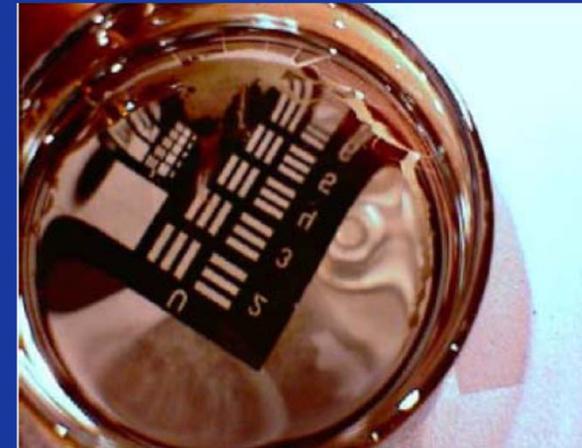
Harley Thronson, NASA

Lunar Telescopes

Roger Angel: Liquid Mirror:

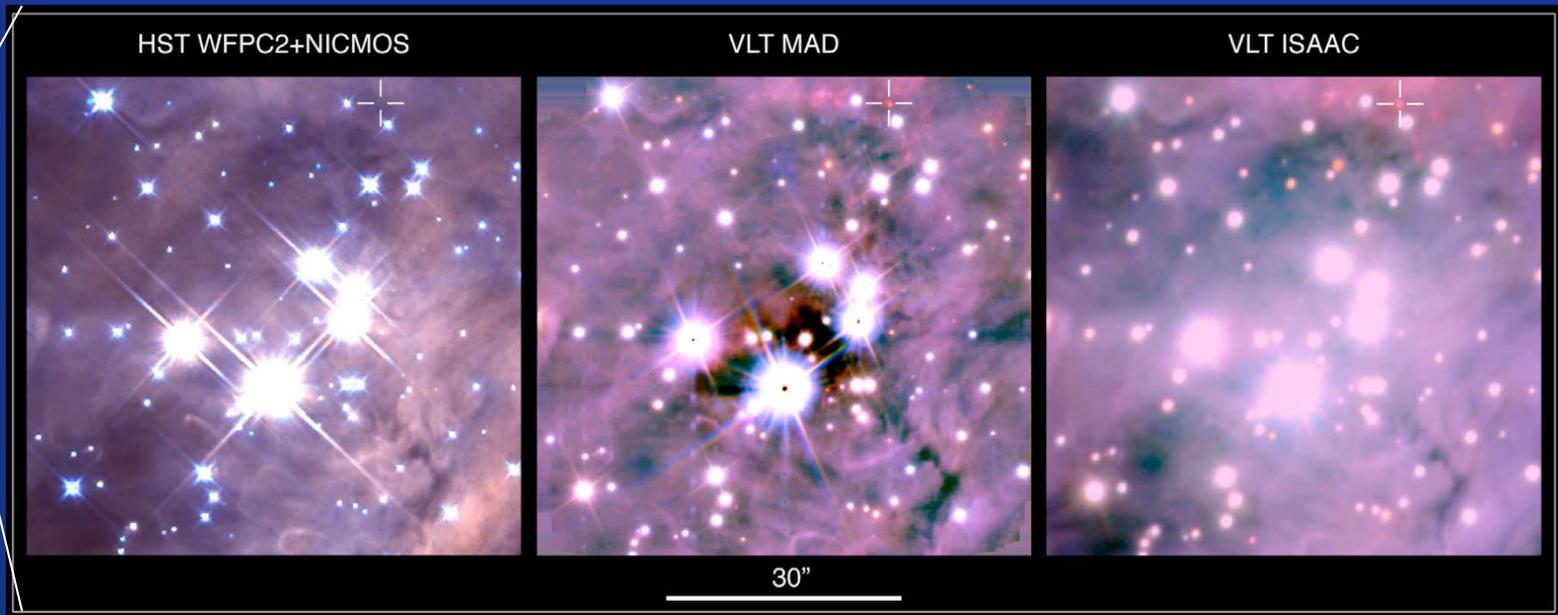
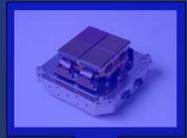
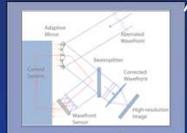
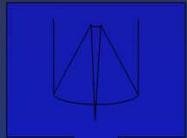
- Metal coated, low viscosity liquid
- Superconducting mag-lev bearing
- Inflating sun-shield

Peter Chen of Goddard has ideas for a spun cast moon dust & carbon nanotube composite mirror!



Roger Angel, Dan Eisenstein, Suresh Sivanandam, Simon P. Worden, Jim Burge, Ermanno Borra, Clément Gosselin, Omar Seddiki, Paul Hickson, Ki Bui Ma, Bernard Foing, Jean-Luc Josset, Simon Thibault, Paul Van Susante

Adaptive Optics MCAO Demonstrator



Hubble Space Telescope

VLT with MAD

VLT with no AO



Origin

PUBLICATIONS OF THE ASTRONOMICAL SOCIETY OF THE PACIFIC

Vol. 65

October 1953

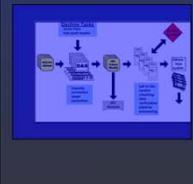
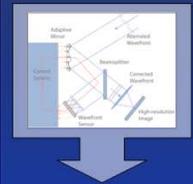
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THE POSSIBILITY OF COMPENSATING ASTRONOMICAL SEEING

H. W. BABCOCK

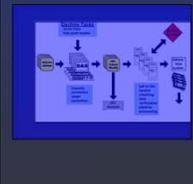
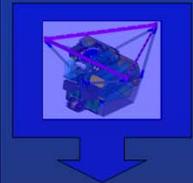
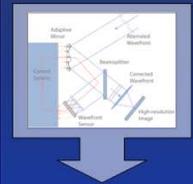
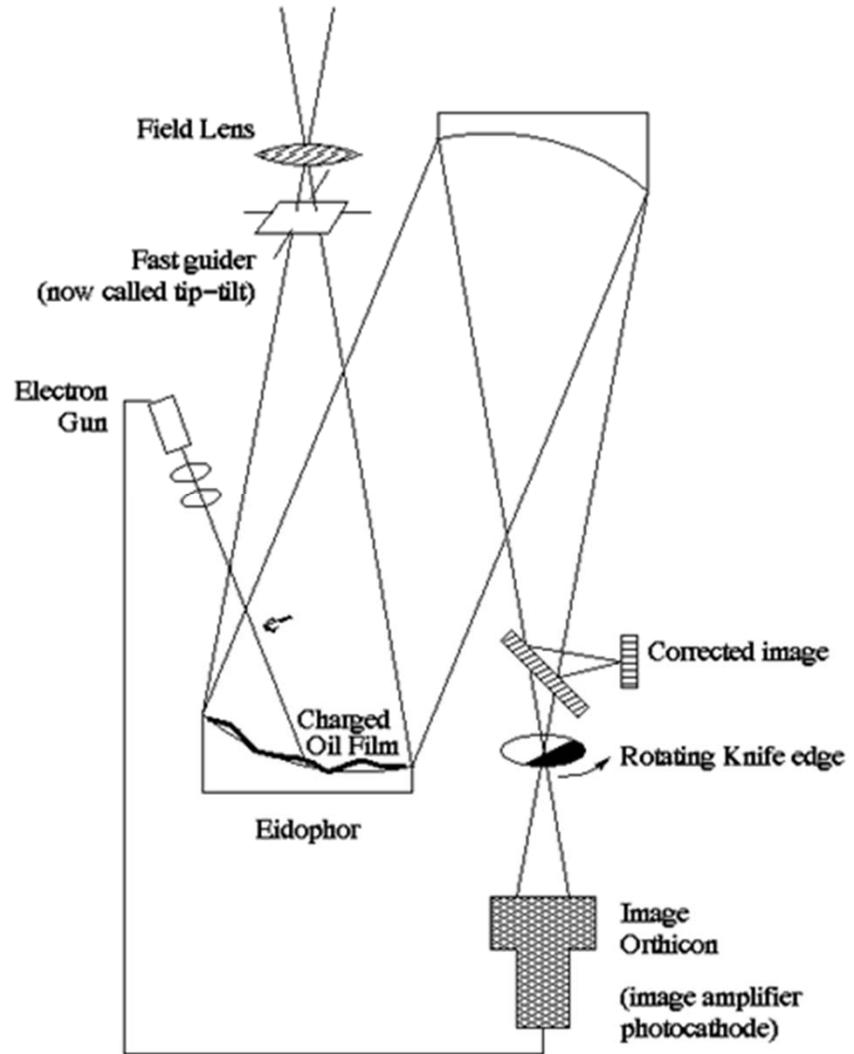
Mount Wilson and Palomar Observatories
Carnegie Institution of Washington
California Institute of Technology

The severe limitations imposed upon nearly all astronomical observations by “seeing”—the effects resulting from passage of light rays through the turbulent atmosphere of the earth—are familiar to every observer. With a small instrument the effect may appear largely as a continual shifting and scintillation of the image of a star, but with a large telescope poor seeing usually





Babcock, H.W., 1953, P.A.S.P. Vol 65, No385
"The Possibility of compensating
Astronomical Seeing"



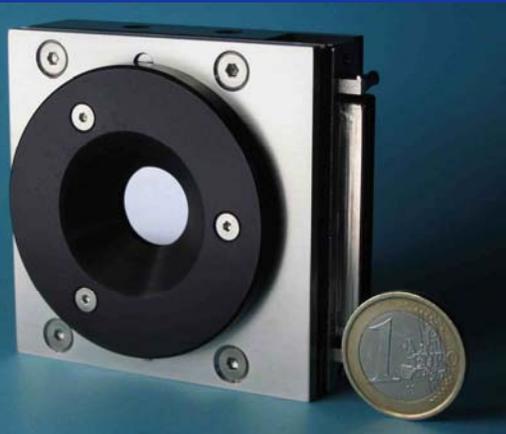
55 Years of Adaptive Optics First Proposed by Horace Babcock

Critical Technology: High density Deformable Mirrors



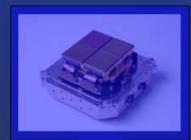
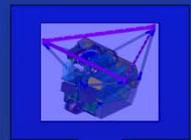
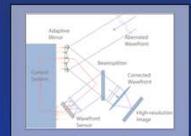
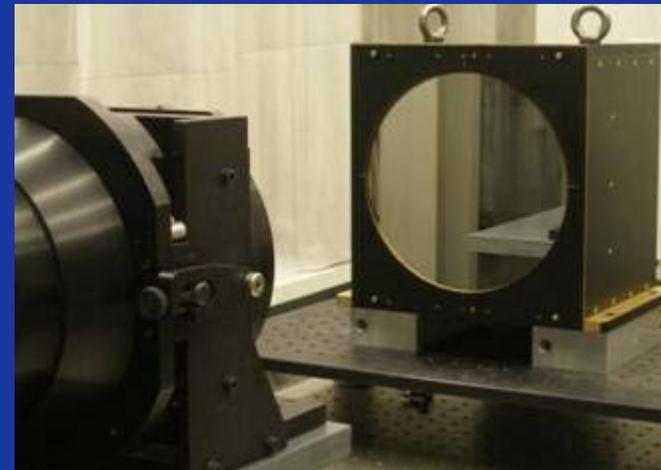
Magnetic DM

- Integration of MEMS and micro-mechanics
- Presently 52 actuators
- Large stroke: up to 100 μm
- Pitch 2mm
- Best flat better than 5 nm
- Aiming to 50 > 100² actuators

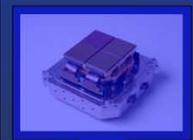
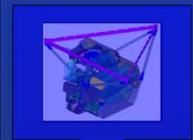
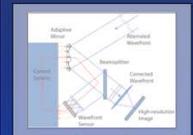


CILAS Piezo Stack DM

- 41x41 = 1370 actuators,
- 4.5 mm pitch
- 9 μm stroke DM

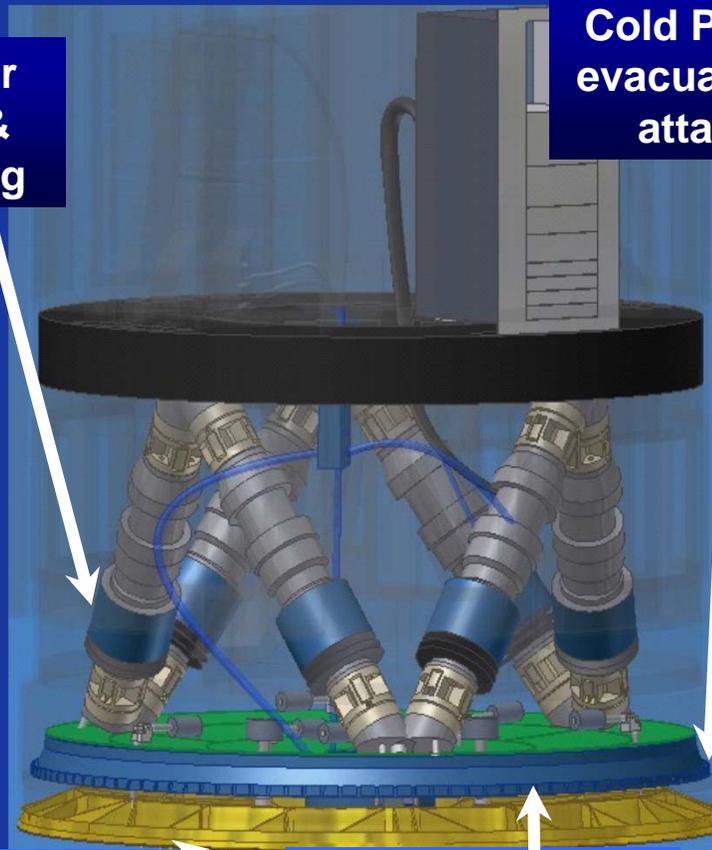


Critical Technology: Large Adaptive Mirrors



Hexapod for centering & fine focusing

Cold Plate; heat evacuation & act. attachment



Thin Shell

Reference body

Voice coil force actuators with capacitive position feedback

New Materials may help!

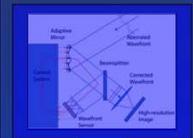
VLT

- 1.1 m Convex
- 1170 actuators
- 29 mm pitch
- 1 ms response
- Stroke <math>< 100 \mu\text{m}</math>

European ELT

- 2.5 m Flat
- 5000 actuators
- 30-50 mm pitch
- 1 ms response
- Stroke 25-90 μm

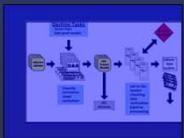
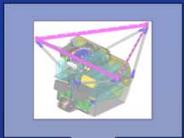
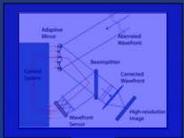
Instruments



VISTA 64 Mpixel IR Camera

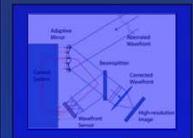
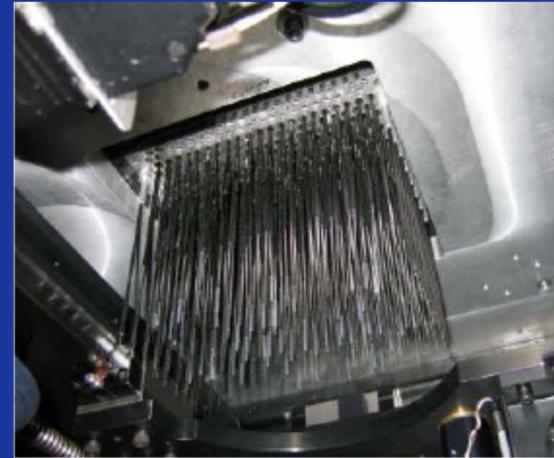
Instruments

- What are the fundamental measurements we do in optical/IR astronomy?
 - Measure intensity and energy of electromagnetic radiation – photons
(And sometimes the amplitude & phase)
 - Divide spectrum up in space or time by dispersion, filtering, Fourier Transform, Fabry-Perot
 - Spatially sample an appropriate field of view
 - **Sounds easy!**



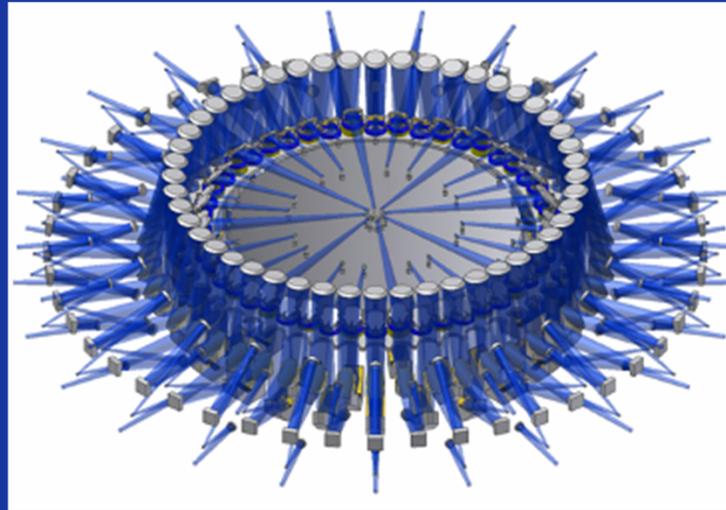
Novel Technologies: Smart Focal Planes

- Make best use of available wide FoV by multiobject and integral field spectroscopy
- Provide alternative to fibre systems for cryogenic infrared instruments

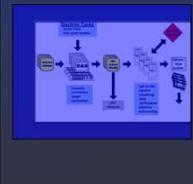
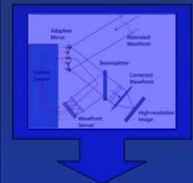


Object Selection

- Options:
 - Pickoffs feeding Integral Field Units and spectrometers
 - Slit exchangers or reconfigurable slits
 - Fibre positioners

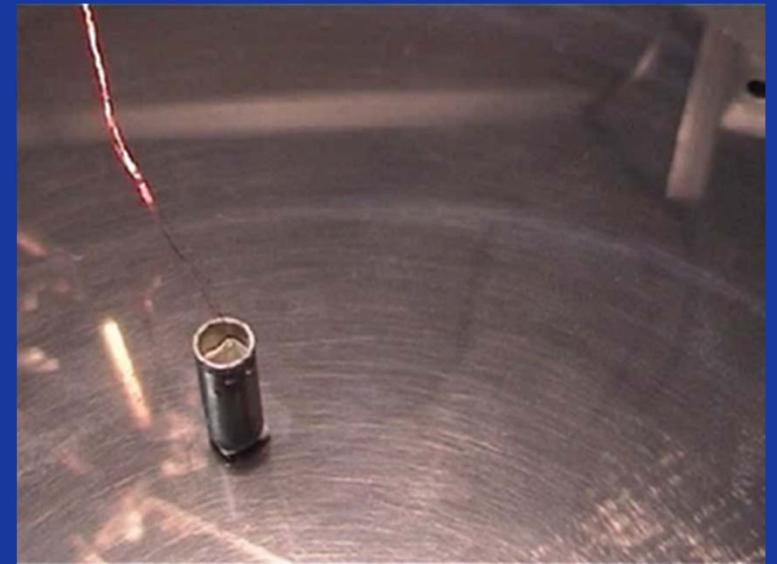
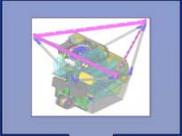
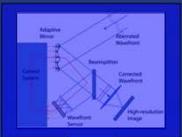


- EAGLE Multi object, multi IFU spectrometer for E-ELT
- Built in Multi Object AO (UK ATC, Durham, LAM, Meudon, ONERA)

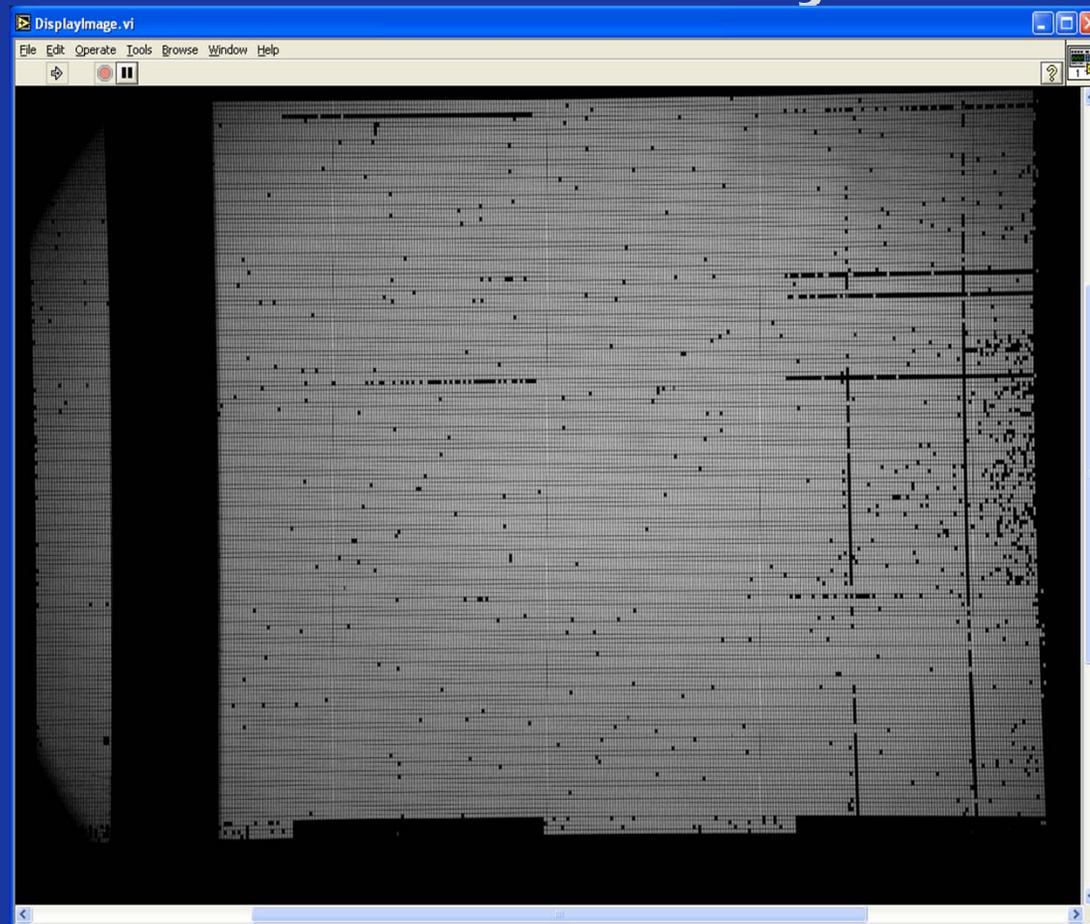
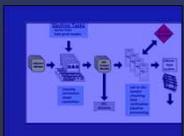
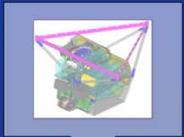
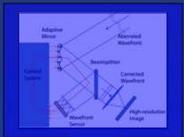




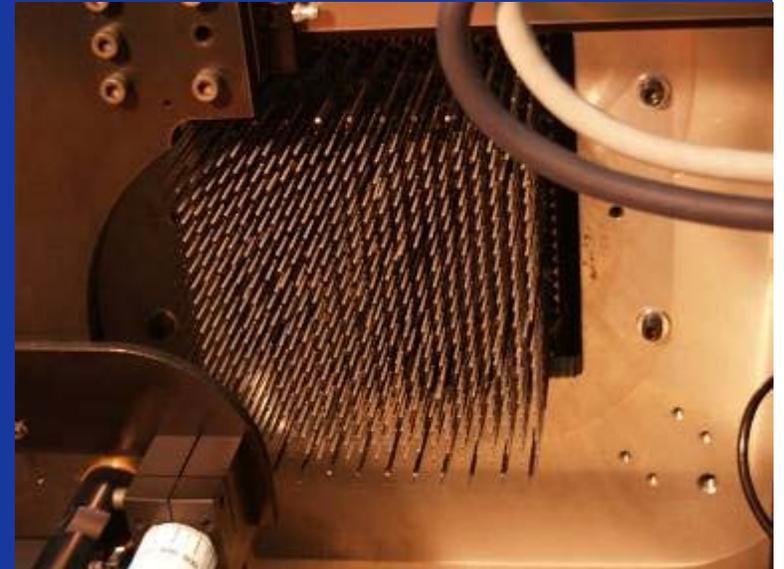
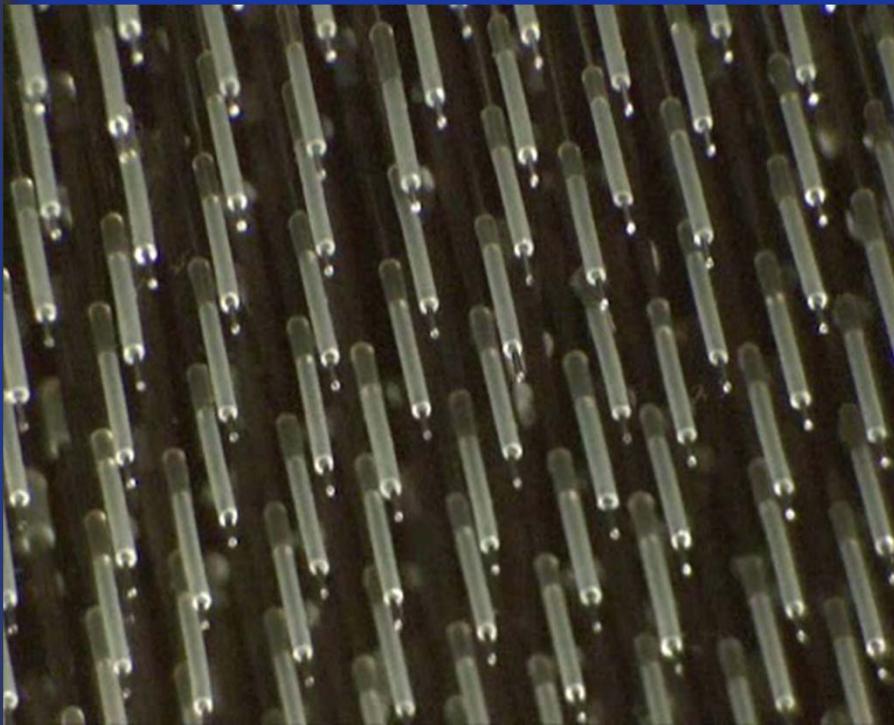
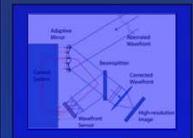
Starbugs & Starpicker



Programmable Slit Spectrometers: MOEMS Shutter arrays



FMOS Echidna: Under testing for SUBARU

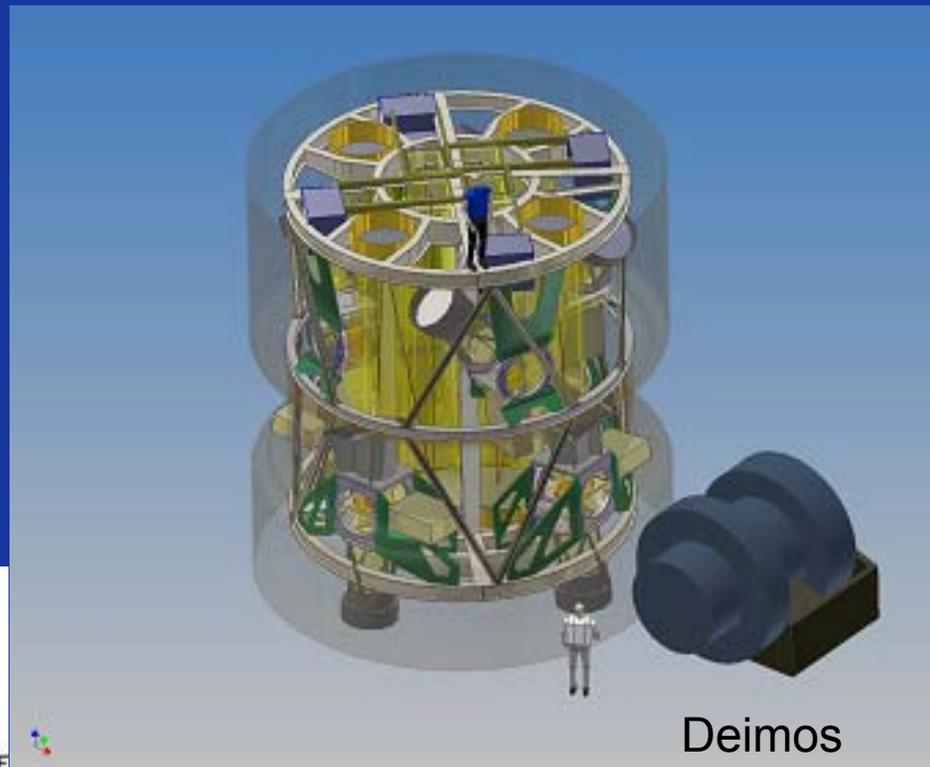
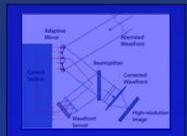
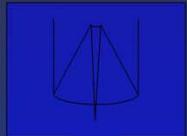


Smart Focal plane with 480 fibres.

Size of seeing-limited ELT instruments!

WFOS Wide Field Optical Spectrograph

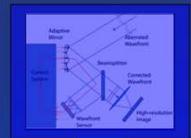
- 8m diam x 10m high
- Size of an 8m telescope!



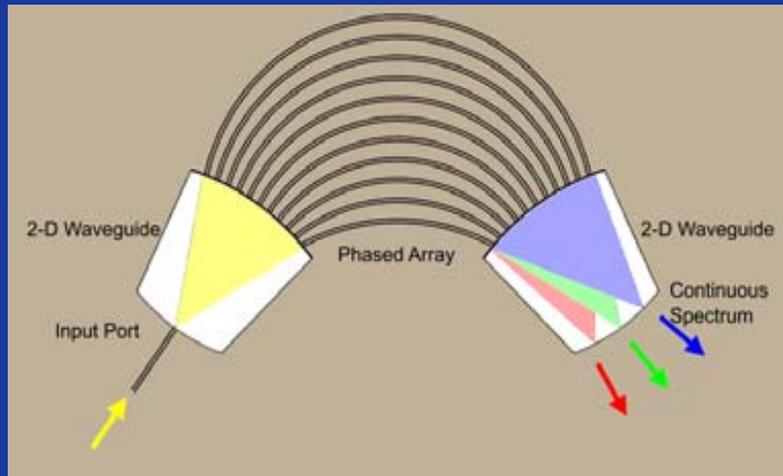
Deimos

Photonic devices

- Devices developed for communications and industrial instrumentation are being investigated for Astronomy:
 - Bragg Gratings
 - Photonic (crystal) fibres
 - Waveguide Beam combiners
- These technologies could be combined with an integrated detector to make compact integrated spectrometers, with built-in OH suppression
- Could we build 1000 spectrometers with integrated detectors to build compact & cheaper multi-object instruments?
- Move them around the focal plane on miniature robots

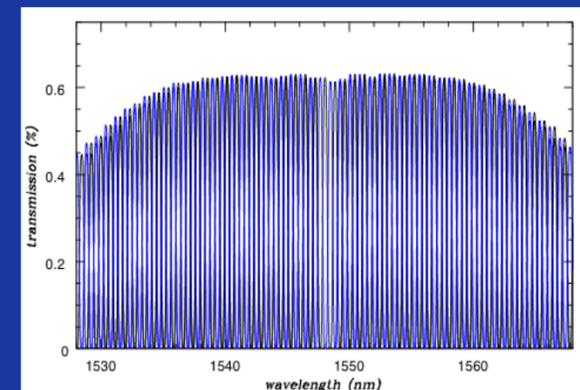


Example: Array Wave Guide devices

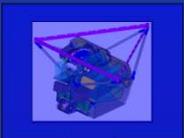
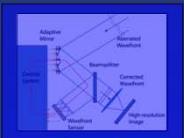
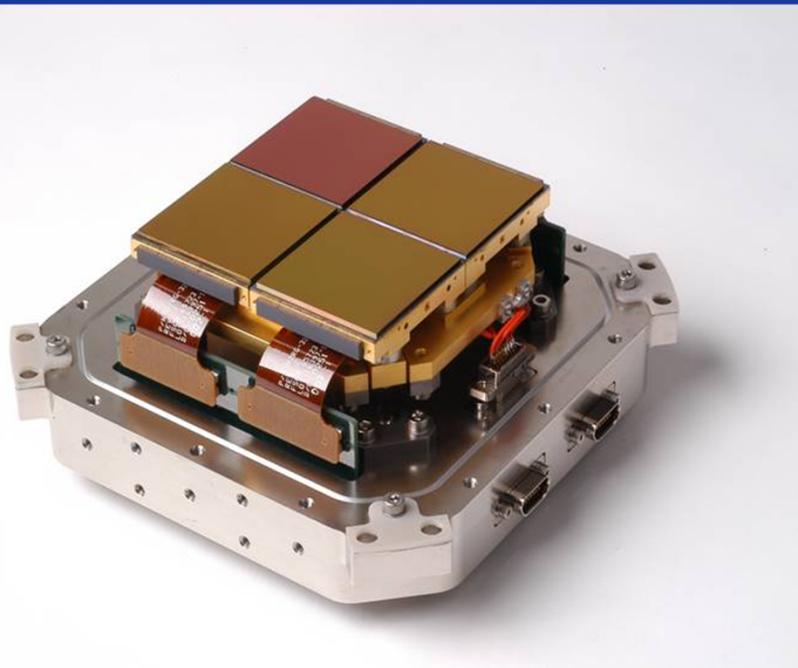


7.5 cm, R=4000

*AAO Newsletter Aug 07
Instruments without optics: an
integrated photonic spectrograph
(Joss Bland-Hawthorn & Anthony
Horton)*

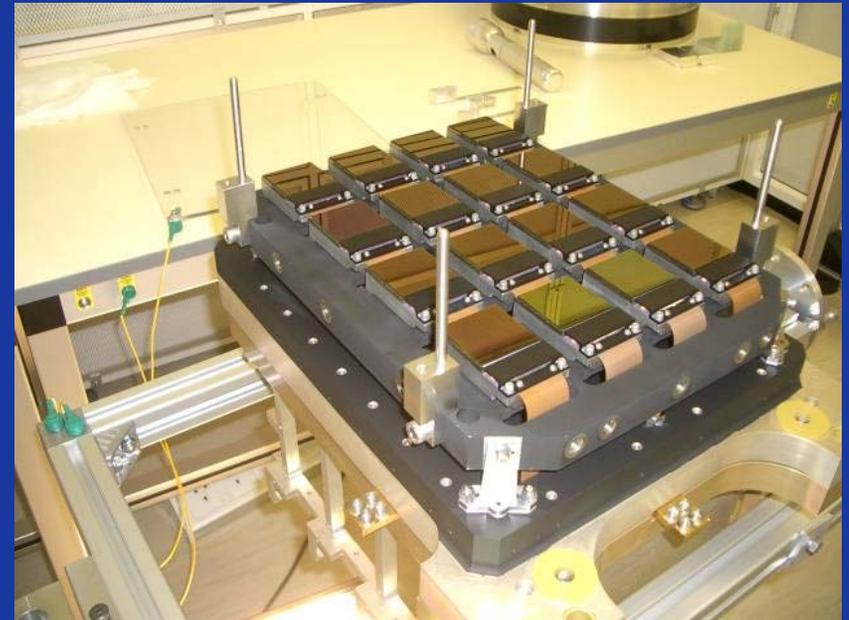


Detectors

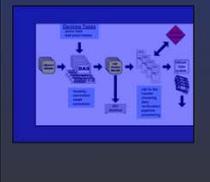
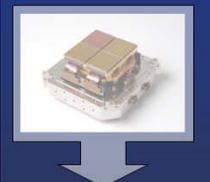
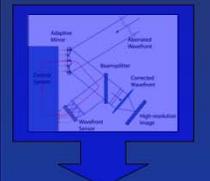
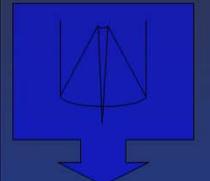


Detectors

- Needs:
 - Bigger, cheaper, better!
- E-ELT will need 60 -100 2kx2k IR arrays
- Lower noise would be good!
- James Webb Space Telescope shows how IR detector performance can be pushed by astronomy
 - But at high cost



16 2kx2k Raytheon Arrays in VISTA

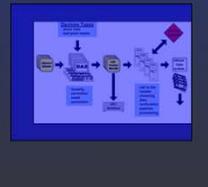
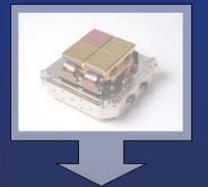
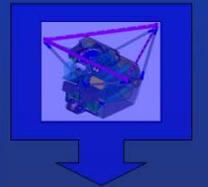
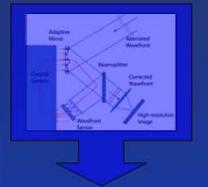


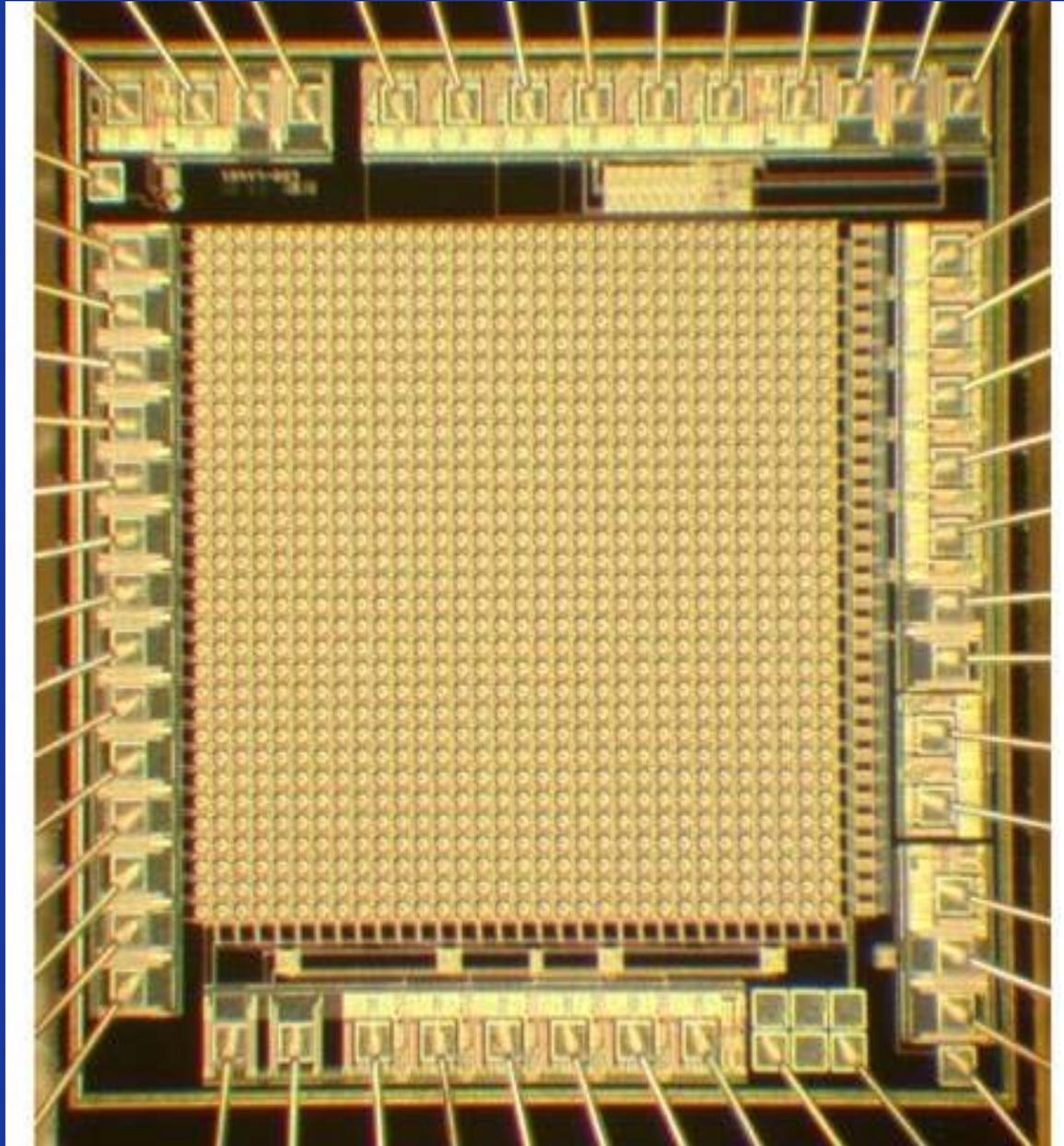
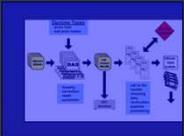
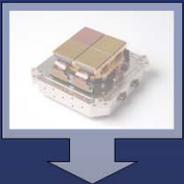
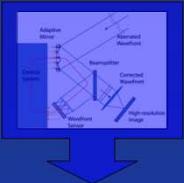
High time resolution

- On timescales of milliseconds
 - study the optical emission from pulsars.
 - optical analogue of the kilohertz quasi-periodic oscillations and related small-scale accretion phenomena found in X-ray binary stars (XRBs)



Accretion processes in an XRB.
Artwork by Catrina Liljegren, *Bild & Form*, Lund; ©Dainis Dravins, Lund Observatory.



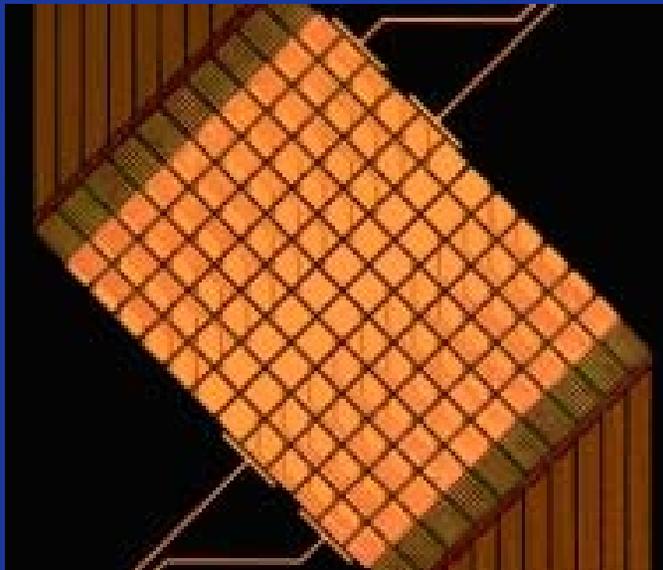


32x32 Single Photon Silicon Avalanche Diode Array
Quantum Architecture Group, *L'Ecole Polytechnique Fédérale de Lausanne*

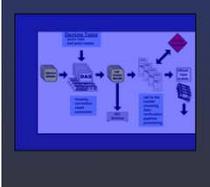
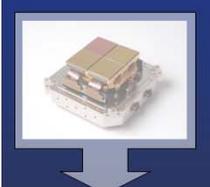
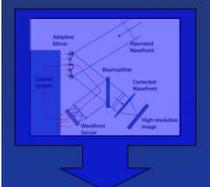
Energy resolving detectors

- Superconducting Tunnel Junctions
- ESA's S-CAM

Detector	12x10 Ta/Al STJs
Pixel size	33 x 33 μm^2
Fill factor	76%
Plate scale	0.8"/pixel
FOV	11"x9"
Pass band	330-745 nm
Maximum detection efficiency (@500nm)	30%
$\lambda/\Delta\lambda$ @ $\lambda=500\text{nm}$	8-11
Event time resolution	$\sim 5 \mu\text{s}$
Operating temperature	285 mK

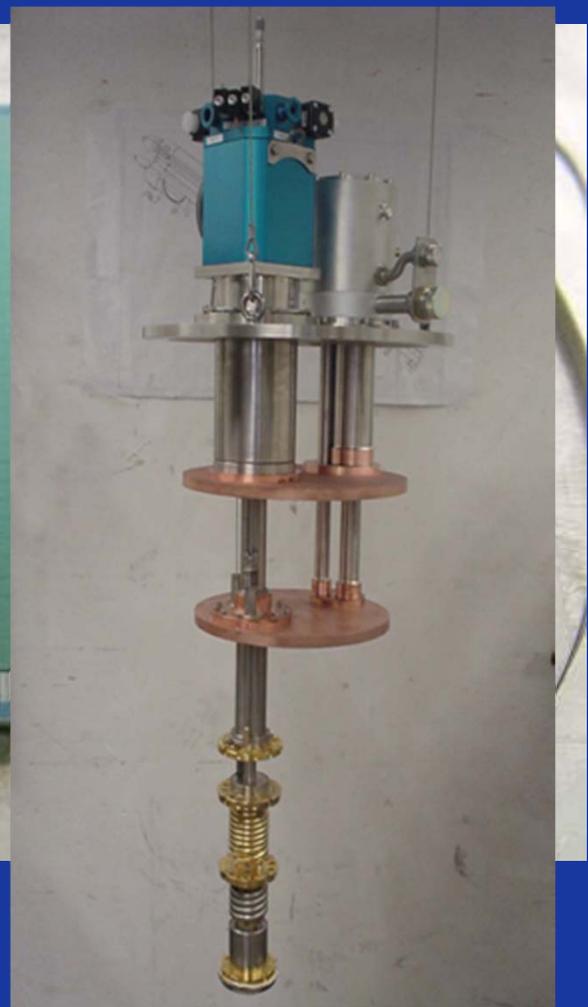
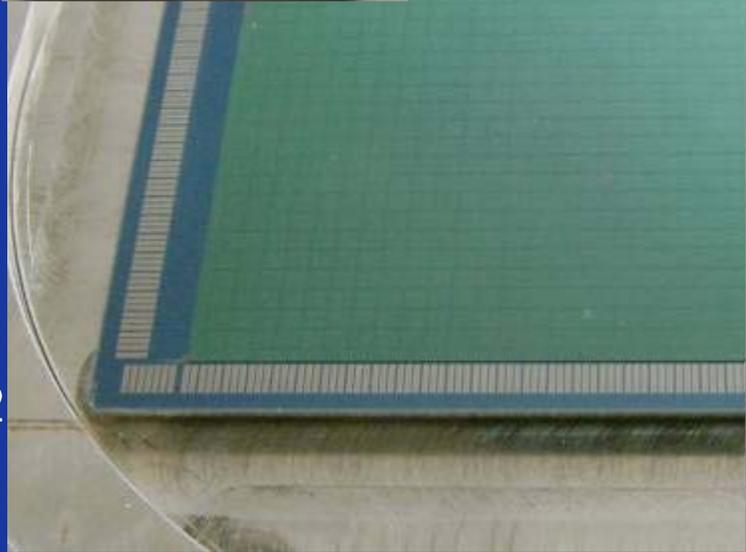
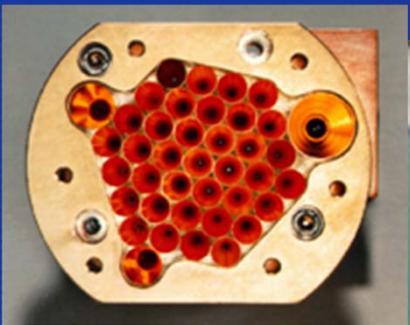


Transition Edge Superconducting Detector Arrays



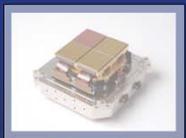
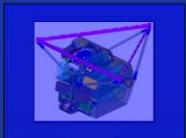
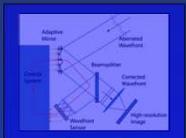
SCUBA
850 μ m array
(same pixel scale)

Completed
40 32
(1280) pixel
array
for SCUBA 2



Other novel detectors....

- Quantum dots
- Quantum Well Interference Photodetector (QWIPs)
- Quantum Well Intrasub-band Photodetector (QWISP)
- Carbon nano-tubes?
 - Bolometers or photodetectors
- Kinetic Inductance Devices
- Single Electron transistors
- Smart Active pixels



Key Technologies for Optical/IR Astronomy

- Next 5-15 years – my view:
 - Formation flying
 - Heavy Launch Vehicles
 - Large and small high-density deformable mirrors
 - Laser Guide Stars
 - Smart Focal Planes
 - OH sky line suppression devices
 - Fast detectors
 - Lower noise (and cheaper!) NIR detectors
 - Free-form aspheric optics
 - Lightweight mirrors
 - Photonic Fibres
 - Laser Comb calibration sources
 - IR optimised fibres
 - Photonic beam combiners
- What are the priorities & what can we afford to develop?
 - Should be driven by science & facility needs
 - And what we can get from developments outside astronomy

What will the Disruptive Technologies of the future be?



By definition – impossible to answer!

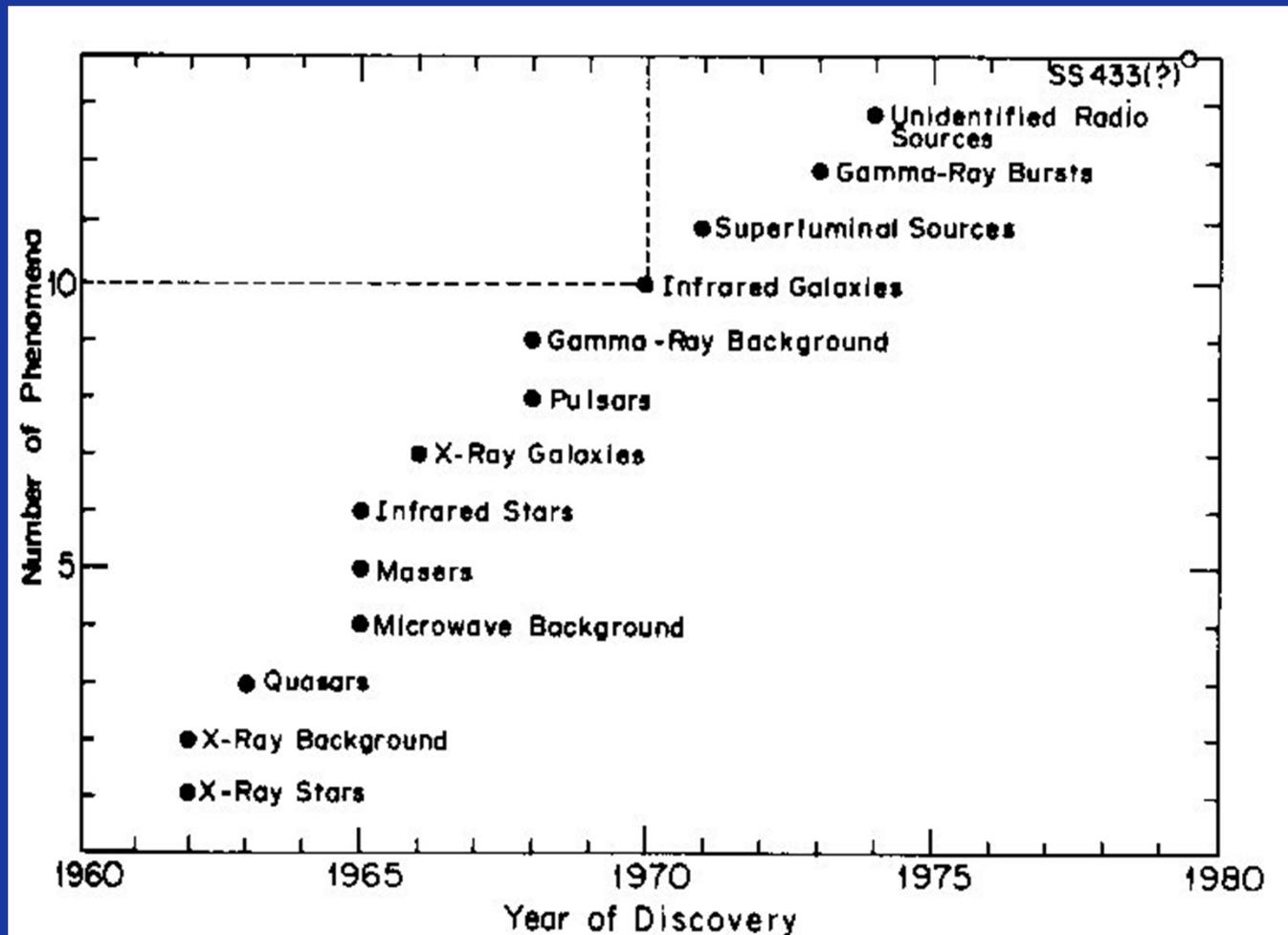
Let's try anyway!

- Photonics:
 - OH Suppression Devices – big improvements in ground-based NIR sensitivity
 - Swarms of Robotic Micro-Spectrometers
- Energy Sensitive Detectors
 - No dispersion elements
- Robotic Assembly of Space Telescopes
 - Very large aperture
 - Interferometers

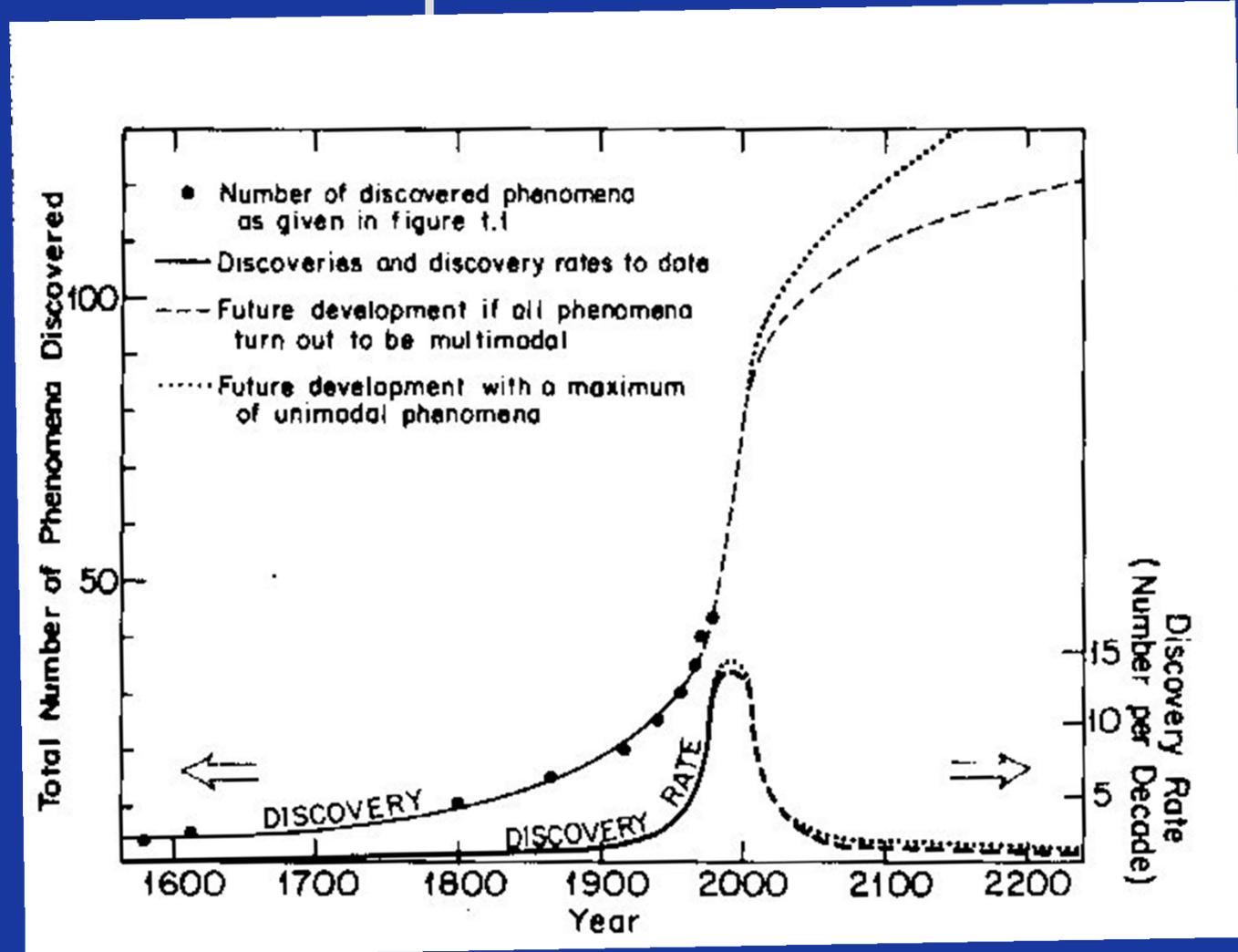
Where will the new parameter space be, when we have JWST, ELTs & Space interferometers?

- Will it be:
 - More sensitivity
 - Bigger, colder telescopes
 - Higher spatial resolution
 - Interferometry
 - Larger Aperture
 - Better adaptive optics
 - High contrast imaging & spectroscopy
 - High time resolution
 - Wider field...
- Where will the balance be: Space/Ground?
- 'Wide Field Surveys and Astronomical Discovery Space' Andy Lawrence (Astronomy & Geophysics, 2007)
 - *Most parameter space is full, or too expensive to extend, except high time resolution and neutrinos!*
- See Martin Harwit's *Cosmic Discovery* from 1981

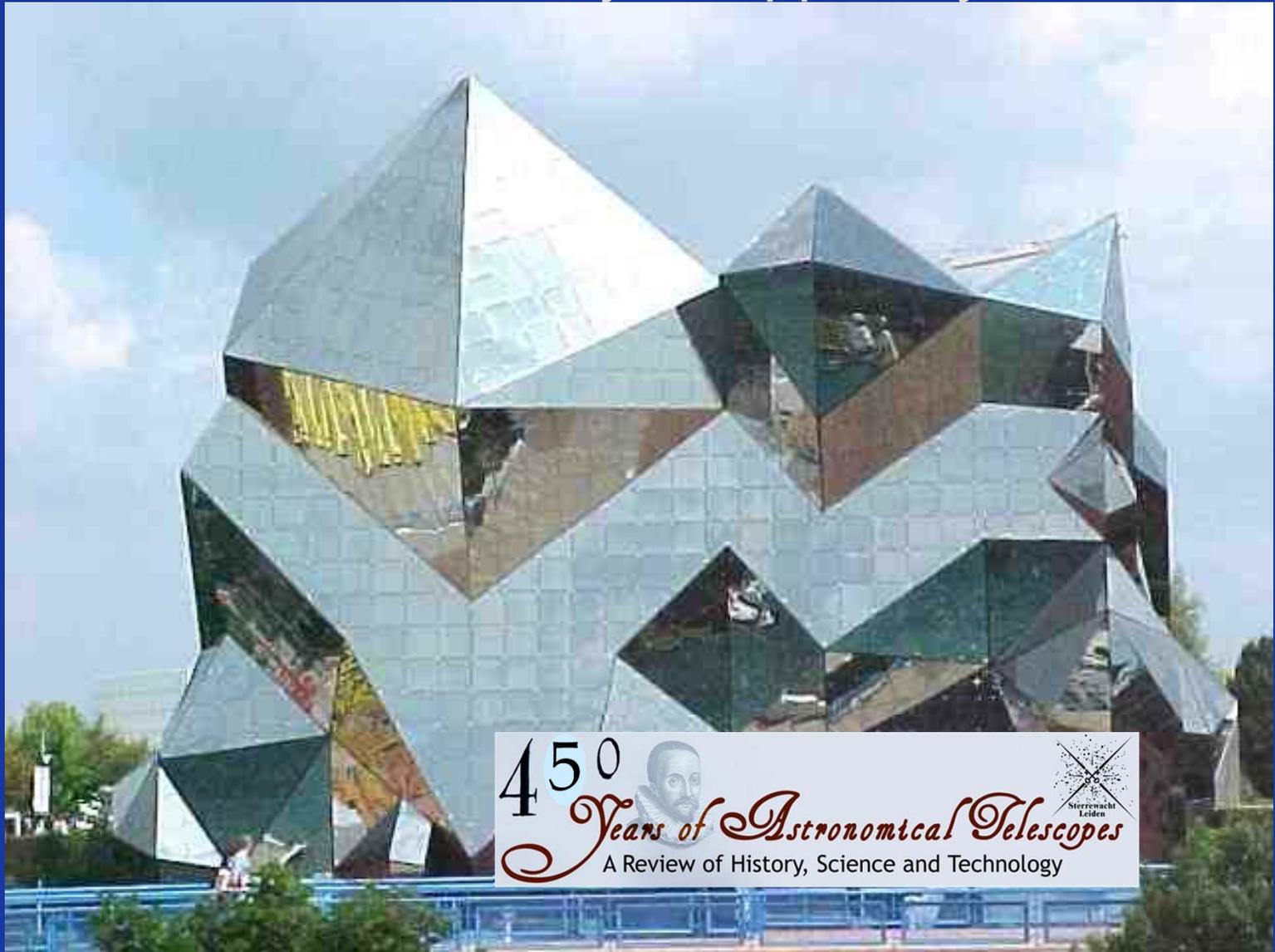
Martin Harwit's prediction in 1981



Martin Harwit's discovery rate prediction



Where will Astronomy be in 50 years: 450th anniversary of Lipperhey?



450
Years of Astronomical Telescopes
A Review of History, Science and Technology



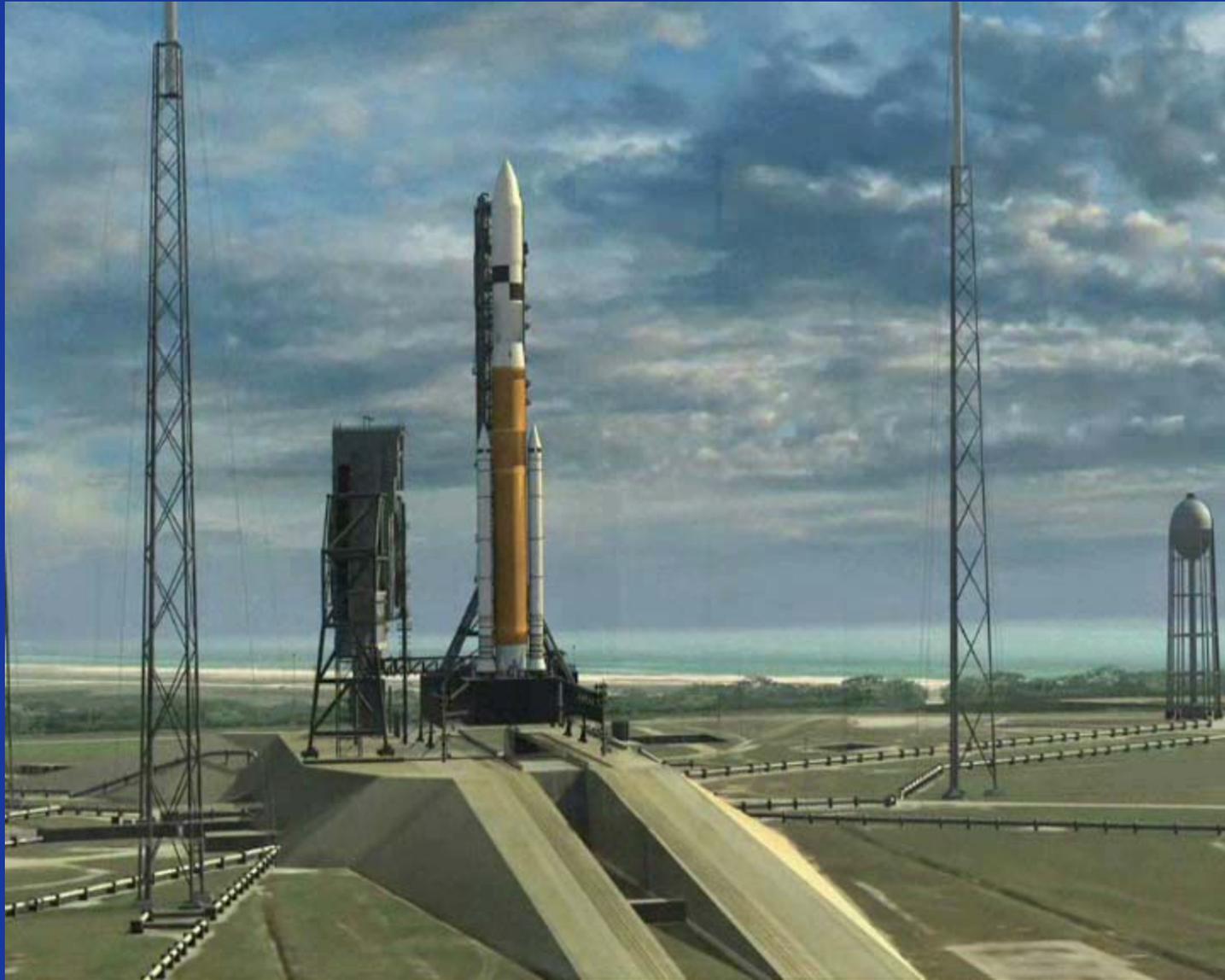
Astronomical Telescopes and Instruments in 2058 ?

- Will we have run out of astronomical discoveries to be made at affordable cost?
 - Is there a law of diminishing returns?
- Or will we have:
 - Nano engineered membrane telescopes
 - Swarms of robotic integrated photometers
 - Interferometers with integrated photonic instruments using detectors measuring intensity, energy and polarisation?
 - Or a liquid mirror telescope on the moon?
- Or a larger version of what we are planning now?
 - ELTs up to 100m – OWL-like?
 - On a cold/high site – Antarctica?
 - With instruments like today's ?
 - 15 – 25 m segmented space telescopes
- Radio Arrays take over – SKA ?
- Or will discovery space not be from Electromagnetic Wave detection but with Gravitational Wave or Neutrino observatories?

Economic and Resource Impacts

- Will the Golden Age of Astronomy be over – due to resource crises, global warming, financial meltdown?
- Will governments, the public and benefactors run out of patience with investment in bigger and bigger facilities?
 - Are we already seeing such a fatigue in space missions – and particle physics?
- But what about the ambitions of China and India?
 - at present these are mostly aimed at non-astronomy space projects

Ares V and 8m Monolithic Telescope



Harley Thronson, and NASA Future In-Space Operations working group